

CHEMICAL CLASSIFICATION OF SPHERULES RECOVERED FROM THE PACIFIC OCEAN SITE OF THE CNEOS 2014-01-08 (IM1) BOLIDE.

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Abstract: We have conducted an extensive towed-magnetic-sled survey during the period of June 14-28, 2023, over the seafloor about 85 km north of Manus Island, Papua New Guinea, centered around the calculated path of the bolide CNEOS 2014-01-08 (IM1). We found about 850 spherules of diameter 0.1-1.3 millimeters in our samples. The samples were analyzed by micro-XRF, Electron Probe Microanalyzer and ICP Mass spectrometry. Here we report major and trace element compositions of the samples and classify spherules based on that analysis. We identified 78% of the spherules as primitive, in that their compositions have not been affected by planetary differentiation. We divided these into four groups corresponding to previously described cosmic spherule types. The remaining 22% appear to all reflect planetary igneous differentiation and are all different from previously described spherules. We call them D-type spherules. A portion of the D-spherules show an excess of Be, La and U, by up to three orders of magnitude relative to the solar system standard of CI chondrites. Detailed mass spectroscopy of 12 of these “BeLaU”-type spherules, the population of which may constitute up to ~10% of our entire collected sample, suggests that they are derived from material formed by planetary igneous fractionation. Their chemical composition is unlike any known solar system material. We compare these compositions to known differentiated bodies in the solar system and find them similar to evolved planetary materials - with lunar KREEP the closest in terms of its trace element enrichment pattern, but unusual in terms of their elevated CI-normalized incompatible elements. The “BeLaU”-type spherules reflect a highly differentiated, extremely evolved composition of an unknown source.

Introduction

The retrieval of cosmic spherules from meteor sites has a long history, with related morphology and composition analyses linking them to various components of the solar system (Brownlee et al., 1979; Maurette et al., 1991; Taylor and Brownlee, 1991; Xue et al., 1994; Brownlee et al., 1997; Herzog et al., 1999; Taylor et al., 2000; Engrand et al., 2005; Genge et al., 2008; Vondrak et al., 2008; Wittke et al., 2013; Folco et al., 2015; Rudraswami et al., 2015; Genge et al., 2017). The spherules range in diameter from a minimum recoverable size of 0.25 to 1.7 mm. They have been classified as I, S and G-types.

Marvin and Einaudi (1967) and others (Folco and Cordier 2015; Blanchard et al. 1980; Brownlee et al. 1997) have listed possible extraterrestrial and terrestrial origins for magnetic spherules like those discussed in this paper. Extraterrestrial origins may be: (i) ablation of meteorites during flight in the atmosphere; (ii) disintegration of carbonaceous chondrites in flight through the atmosphere; (iii) vaporization of large meteorites during impact crater formation; and (iv) infall of particles containing iron or iron oxide, of either asteroid or comet origin (they may enter the atmosphere as spherules, resulting from collisions in space, or may become spherules in the atmosphere). The terrestrial origins proposed were: (i) volcanic; (ii) industrial combustion of coal, crude oil, or wood; (iii) smelting products; and (iv) others such as forest fires and lightning discharges.

On 8 January 2014 US government satellite sensors detected three atmospheric detonations in rapid succession about 84 km north of Manus Island, outside the territorial waters of Papua New Guinea (20 km). Analysis of the trajectory suggested an interstellar origin of the causative object CNEOS 2014-01-08: an arrival velocity relative to Earth more than $\sim 45 \text{ km s}^{-1}$, and a vector tracked back to outside the plane of the ecliptic (Siraj and Loeb 2022a). In 2022 the US Space Command issued a formal letter to NASA certifying a 99.999% likelihood that the object was interstellar in origin. Along with this letter, the US Government released the fireball light curve as measured by satellites, which showed three flares separated by a tenth of a second from each other. The bolide broke apart at an unusually low altitude of $\sim 17 \text{ km}$. The object was likely substantially stronger than any of the other 272 objects in the CNEOS catalog, including the $\sim 5\%$ -fraction of iron meteorites from the solar system (Siraj and Loeb 2022b). Calculations of the fireball light energy suggest that about 500 kg of material was ablated by the fireball and converted into ablation spherules with a small efficiency. The fireball path was localized also based on the delay in arrival time of the direct and reflected sound waves to a seismometer located on Manus Island (Siraj and Loeb 2023), but this inference has been debated (e.g., Brown and Borovicka 2023). In this paper, we characterize spherules retrieved in an expedition that surveyed the region identified by the US Government satellites as the meteor site without assuming their association with the IM1 bolide.

Sampling Expedition

The expedition was mounted from Port Moresby, Papua New Guinea (PNG), to search for remnants of the bolide, labeled hereafter IM1. It utilized a 40-meter catamaran workboat, the M/V Silver Star. A 200-kg sled (**Fig. 1**) was used with 300 neodymium magnets mounted on both of its sides and video cameras mounted on the tow-bridle. Approximately 0.06 km^2 were sampled in the target area (**Fig. 2**). The fine material collected on the neodymium magnets was extracted and brought in a wet slurry up to a laboratory set up on the bridge of the vessel for further examination. There, an initial wet-magnetic separation took place. Subsequently, both magnetic and non-magnetic separations were processed through sieves and dried. Spherules were handpicked with tweezers using a binocular zoom microscope. They ranged in size from 100 microns to 2 mm. We obtained a total of ~ 850 fragments consisting of spherules and shards by this method.

Spherule Samples and Location

The terminology and measurements will be explained in the upcoming subsections (see also [Loeb et al. 2024a,b](#)). Figure 2 shows the tracks of our expedition survey. Tracks 22 and 17 were located outside the Department of Defense error box for IM1's fireball, but the full extent of IM1's strewn field is unknown.



Figure 1: Magnetic sled design. Left: Isometric view of the magnetic sled design, Right: Side view of the sled when placed on the ocean floor. The 250-kg, 1 by 2 meter sled was covered with an array of 300 neodymium magnets on both sides and equipped with video cameras in the metal tow-halter ahead of it, which was anchored by a synthetic cable to a winch on the ship, the M/V Silver Star.

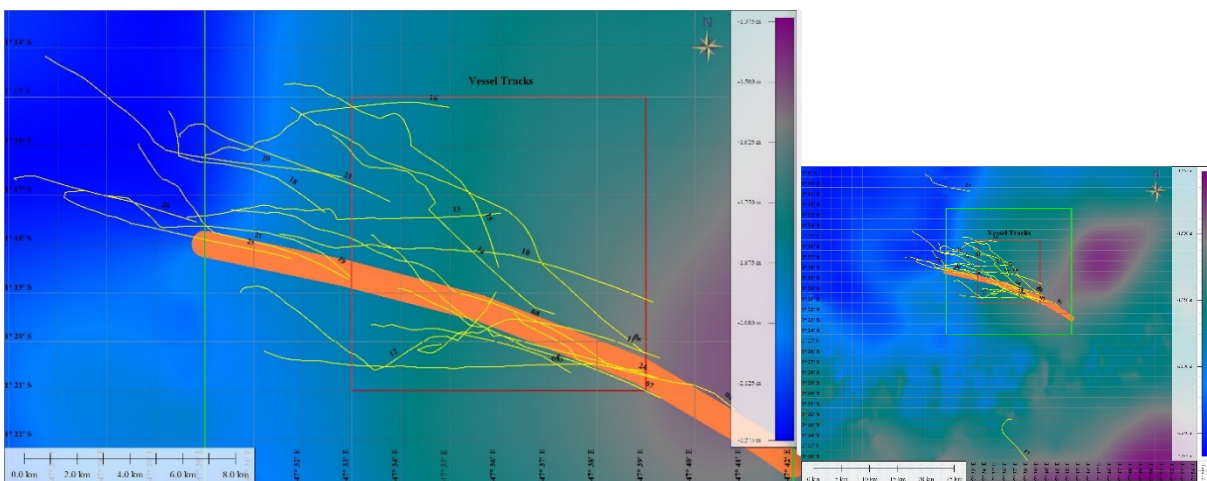


Figure 2: List of ship track numbers around the expected path of IM1 based on seismometer data ([Siraj and Loeb, 2023](#)) from Manus Island (orange strip). Left: Tracks within the Department of Defense error box (in red). Right: Full map including runs 22 and 17. Background colors indicate ocean depth (with scale on the right). Latitude and longitude are marked in degrees and decimal minutes. At one degree south latitude, one minute of latitude or longitude equals one nautical mile which is 1.852 km. The red box, measuring 11.112 km on a side, marks the uncertainty in the Department of Defense (DoD) localization of IM1's fireball, and the green box marks twice that size.

Analytical Methods

Most samples (~802) were first analyzed by microXRF with a Bruker Tornado M4 for their bulk major element composition, followed by imaging (SEM and EDS chemical mapping) and spot chemical analyses of ~80 samples with a JEOL Model JXA 8230 Electron Probe Microanalyzer (EPMA). Measurements of elemental abundances for about 60 major and trace elements were performed for 68 samples with an iCAP TQ triple quadrupole ICP-MS (ThermoFisher Scientific).

microXRF: The major and trace element compositions of separated spherule and shard samples was analyzed using the M4 TORNADO PLUS micro XRF instrument from Bruker, situated at the central application facility in Berlin, Germany. This instrument features a 30 W micro-focus X-ray tube and a rhodium anode, with X-rays focused through a polycapillary lens to achieve a spot size of 20 μm for the high-energy range above 20 keV. The measurements were conducted under specific conditions: 50 kV, 200 μA , and 2 mbar for a duration of 60 s. The pressure of 2 mbar was chosen to enhance the detection of light elements and minimize absorption caused by air between the sample and detector. Fluorescent signals were collected using two light element window silicon drift detectors, each with a 60 mm^2 active area, with

independently operating signal processing units, allowing a maximal throughput of 250 kcps each at a spectroscopic resolution lower than 145 eV for manganese K α . The samples were analyzed using the single-spot approach, capitalizing on the benefits arising from the significant information depth of X-rays (e.g. Beckhoff, B. et al., 2006) and the beam divergence of the poly-capillary lenses (approximately 60 μ m/mm). This combination ensured optimal coverage of the sample volume and yielded representative results. Sample quantification was attained through standard-less fundamental parameter quantification after instrument calibration, involving a mix of pure elements and the NIST 620 certified reference glass sample to describe lens transmission and element sensitivity. The quantification algorithm uses the forward calculation of the complete spectrum following the approach describe by Sherman (1955). To eliminate any potential for blank signals, spherules and shards were mounted on acrylic plates using double-sided tape. The results for major and trace elements were computed, presenting major elements as oxides and traces as individual elements. The primary advantage of micro XRF lies in the combination of speed, major and trace element sensitivity as well as its non-invasiveness, positioning it as a pre-screening tool for more time-consuming and invasive analytical procedures with uncertainties below tens of percent.

EPMA: We used the JEOL JXA-8230 electron microprobe at the Harvard Electron Microprobe Laboratory to obtain high-resolution backscatter (BSE) and secondary (SEI) electron images, elemental X-ray maps, and chemical analyses, using focused beam of $\sim$ 1 mm in diameter. Most objects were mounted on sticky tape, while several spherules were mounted in epoxy and polished. All samples were carbon-coated. The chemical compositions of intact objects were measured by energy dispersive spectroscopy (EDS) using factory calibration curves. The polished areas of several spherules were analyzed by wavelength dispersive spectroscopy (WDS) using common natural minerals and synthetic glasses as calibration standards (e.g., Petaev and Jacobsen, 2009). The EDS analyses and imaging we performed at accelerating voltage of 20 kV and beam currents of 5-10 nA and $\sim$ 0.1 nA at low- ($< 1000\times$) and high-resolution ($> 2000\times$), respectively. In the WDS analyses we used accelerating voltage of 15 kV, beam current of 20 nA, and counting times of 30 sec and 15 sec on peak and background, respectively.

TQ-ICP-MS (Triple Quad elemental analysis): Measurements of elemental abundances for major and trace elements were performed on the iCAP TQ quadrupole ICP-MS (ThermoFisher Scientific) in the Cosmochemistry Laboratory at Harvard University. USGS reference materials were thoroughly dissolved and diluted in a 2% HNO₃ solution spiked with 10 ppb indium diluted to a factor of 5000 to be used as standards. Spherules were prepared for mass spectrometry measurements by first individually digesting the samples in a mixture of concentrated HF-HNO₃-HCl at a 1:3:1 ratio at 120-140 $^{\circ}$ C overnight. The samples were subsequently dried down and then redissolved in a second acid mixture involving an aqua regia solution mixed with H₂O at a 3:2 ratio and heated to 120-140 $^{\circ}$ C overnight. This dissolution was dried down for the second time and redissolved in a high-purity 2% HNO₃ solution. A small aliquot (3%) was drawn from this solution and further diluted for elemental analysis. To account for and to correct instrumental drift, the 2% HNO₃ solution used for dilution was spiked with 2 ppb indium as an internal standard, prepared identically to the standard solutions. Measurements were performed in KED mode with He as a collision cell gas as recommended by the Reaction Finder function built in the iCAP TQ Qtegra software, with the exception of Cr, which was measured in TQ mode with O₂ as a mass-shifted molecule. The prepared spherule solutions were measured as an unknown against a four-point calibration line consisting of a blank and three USGS standards: BCR-2, BHVO-2, and AGV-2. Calibration curves for individual elements were checked for linear intensity to concentration correlations for accurate measurements. Routine measurements of AGV-2 as an unknown on the iCAP TQ suggest fractional errors to be within 6% using this method.

Results: Imaging and morphology of the spherules

Electron microprobe images of recovered spherules are shown in **Figs. 3, 4 and 5**. The spherules in **Fig. 3** are what we call primitive spherules, identified as such by their chemical composition (see next section). The dendritic textures of these spherules suggest rapid cooling. The spherules in **Fig. 4 and 5** are

all identified as differentiated spherules by their chemical compositions (see next section) and do not in general have as perfect a spherical shape as the primitive group, and some of them are clusters of coalesced spherules. The complete spherule dataset will be released in a future publication. The shapes of the BeLaU-type spherules in Fig. 4c and f have been seen before in spherules from airbursts ([Tankersley et al. \(2024\)](#); [van Ginneken et al. \(2021, 2024\)](#)). In addition to electron microprobe images, low resolution photos were taken for 684 samples prior to micro-XRF analysis. Among these, ~64 were determined to be “shards” as opposed to “spherules,” which were discerned based on sharp angles resembling broken surfaces or evidence of fragmentation.

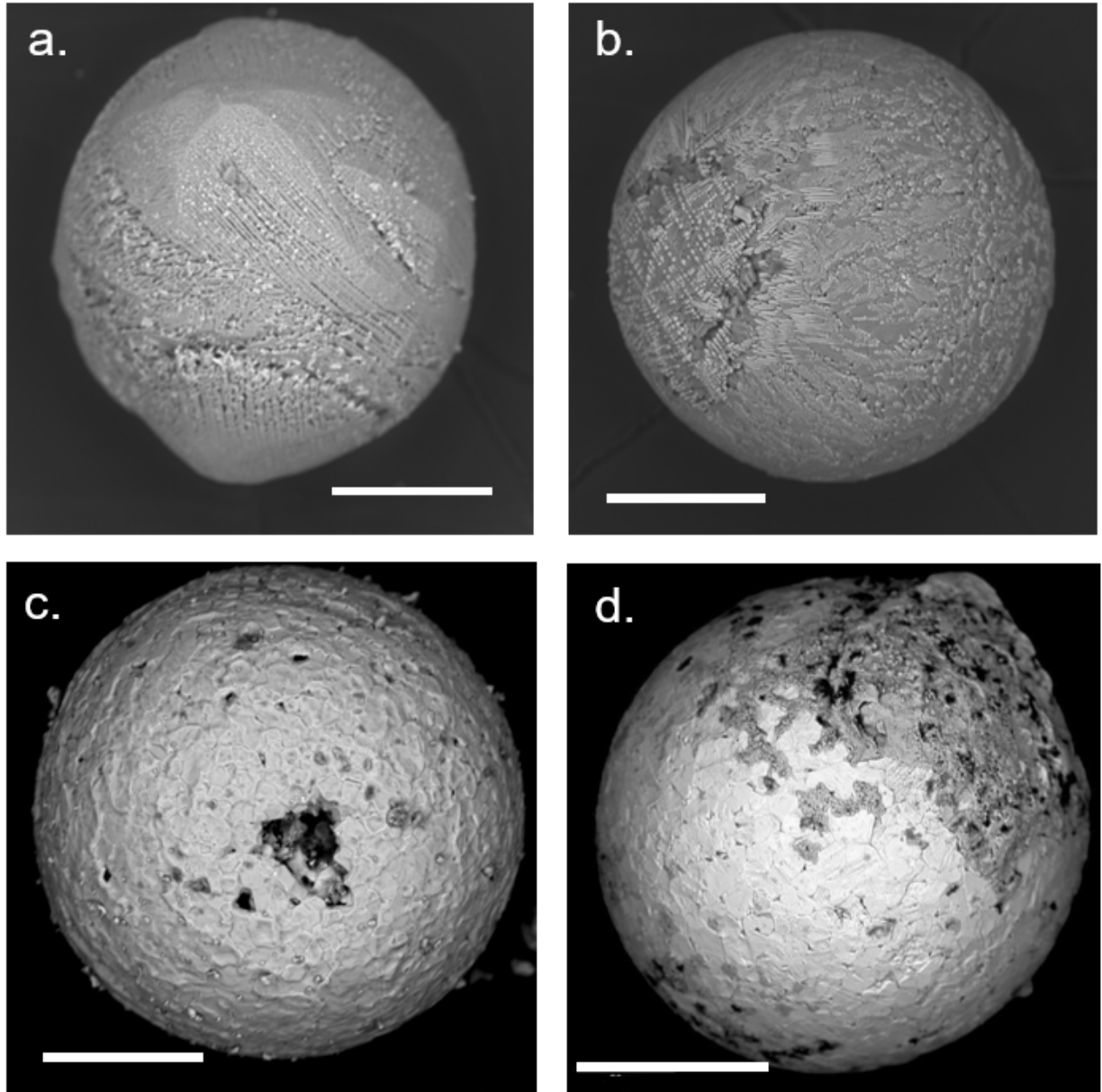


Figure 3. BSE images of primitive spherules from the a) S-type (IS20M-1) b) G-type (IS20M-21) c) I-type, high Ni (IS8M2-20) and d) I-type, low Ni (IS16A-SPH1) groups. The scale bar is 100 microns.

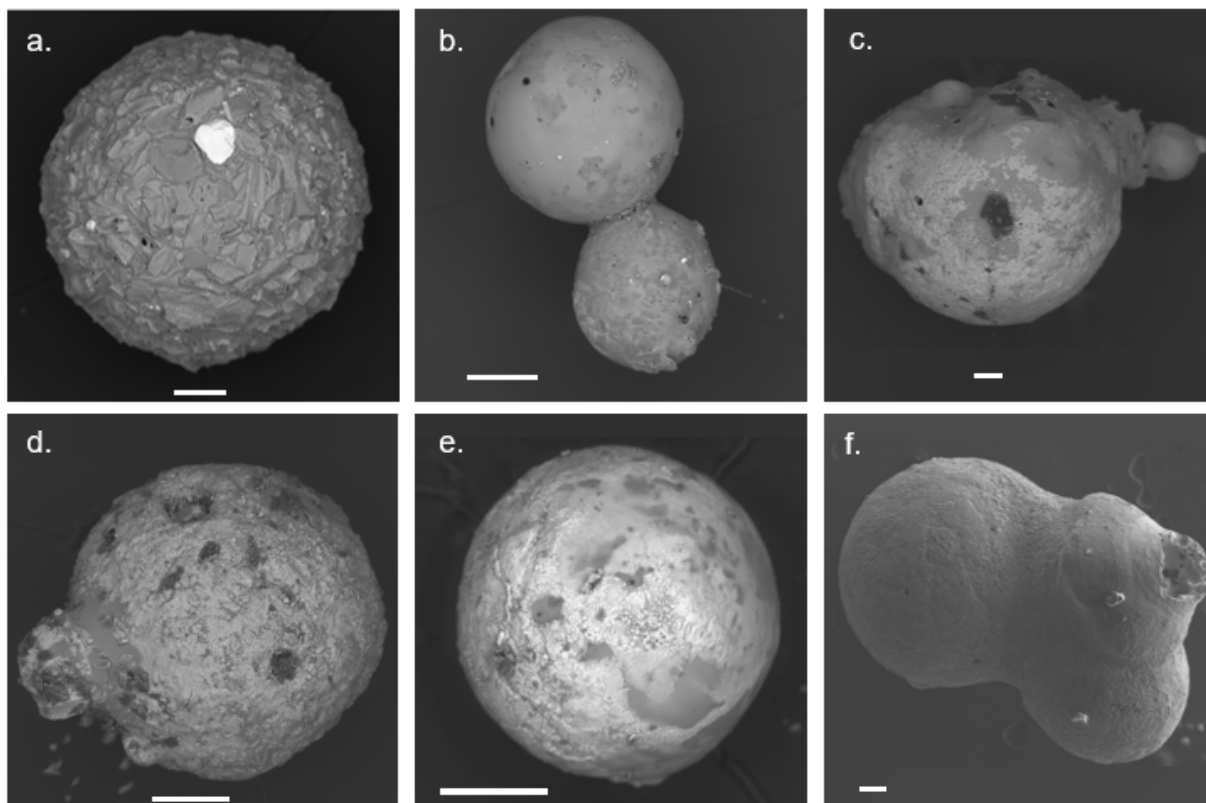


Figure 4. BSE images of differentiated spherules from each of the differentiated (D-type) spherule classifications: a) D-type low Sr, high Si (IS11M2-2) b) D-type high Sr, high Si (IS19M-14) c) D-type, high Si (17NMAG-5) d) D-type, low Sr, low Si (IS21-4) e) D-type high Sr, low Si (IS14M-2) and f) D-type, low Si (S21 or IS14-SPH1) groups. The images c) and f) are also classified as “BeLaU”-type spherules. The scale bar is 100 microns.

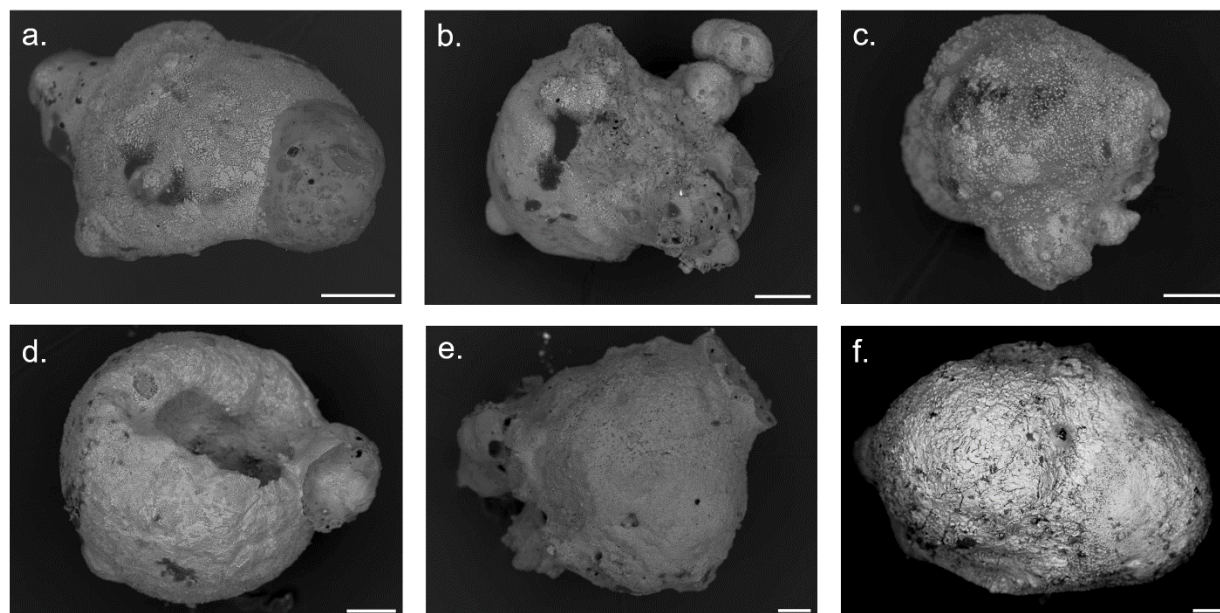


Figure 5. Additional BSE images of D-type spherules: a) 19NMAG-52 (D-type, high Sr, low Si), b) 19MAG-50 (D-type, high Sr, low Si), c) 19MAGx-4 (BeLaU, low Si), d) 17MAG-2 (BeLaU, low Si), e) 17MAG-29 (BeLaU, low Si), and f) 17MAG-6 (D-type, high Sr, low Si). The scale bar is 100 microns.

Classification of the spherules based on elemental compositions.

Cosmic spherules are sub-divided into three compositional types (Blanchard et al., 1980). These are the silicate-rich spherules or S-type, the Fe-rich spherules or I-type and glassy spherules or G-types. There is strong evidence that S-type cosmic spherules with chondritic-like composition are related to carbonaceous chondrites and ordinary chondrites (Blanchard et al., 1980). It has been suggested that I-type spherules are metal grains released from carbonaceous chondrites disaggregated in space (Herzog et al. 1999). The glassy or G-type spherules are thought to be mixtures of S and I. Relatively rare spherules have been called *differentiated* as they have similarities to achondrite meteorites and have been treated as a subgroup of S-type spherules. Differentiated spherules have major-element compositions with higher Si/Mg and Al/Si ratios, and higher refractory lithophile trace element contents relative to chondritic spherules (cf. Folco and Cordier 2015).

The major element, Sr and Ni compositions of 745 spherules from the IM1 site measured by micro-XRF are in **Table S1** (supplement). The data are plotted in a Mg/Si histogram (**Figure 6**) and show a minimum at $\text{Mg/Si} = 1/3$. Spherules with $\text{Mg/Si} > 1/3$ are similar to chondritic meteorites and cosmic S-type spherules, while those with $\text{Mg/Si} < 1/3$ are similar to igneous rocks from Earth and other planetary bodies. The data are also plotted in a Mg-Si-Fe ternary diagram (**Figure 7**), since such a diagram has been shown to effectively distinguish the S-, I- and G-type groups (cf. Folco and Cordier 2015). We note that there are two distinct groups of spherule compositions in this plot. About 78 % of the spherules fall along the trend of S, G and I-type spherules. These are referred to as primitive spherules as they are thought to be related to primitive chondritic meteorites and represent materials that have not gone through planetary differentiation. The remaining 22% of the spherules have low Mg and plot close to the Si-Fe side of the diagram. The high-Si part of this group plots within the range of terrestrial igneous rocks that are shown for comparison. These spherules are thus called differentiated, meaning they are likely derived from crustal rocks of a differentiated planet. Since they are clearly different from the differentiated subgroup of S-type spherules we give them a new name *D-type spherules*. The $\text{Mg/Si} = 1/3$ is used to distinguish primitive and differentiated spherules. The D-type spherules were discovered in 17 of the 24 tracks explored during this expedition.

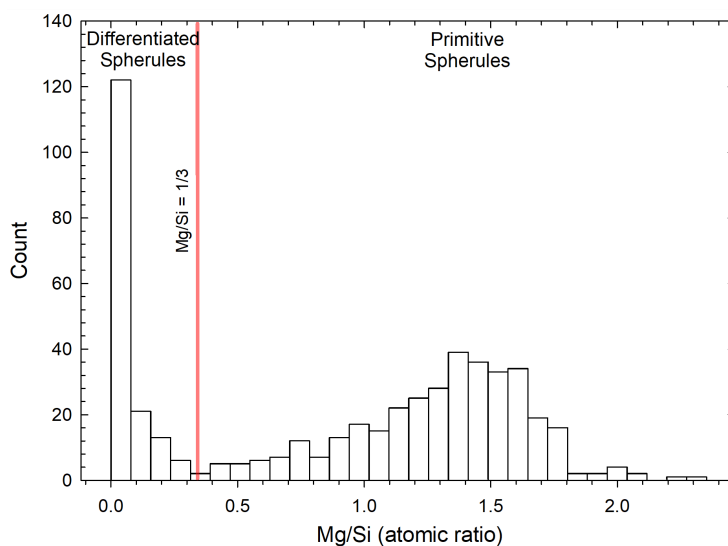


Figure 6. Histogram of the Mg/Si ratio measured with micro-XRF for 745 IM1 site spherules. This diagram shows a clear dividing line between primitive and differentiated spherules.

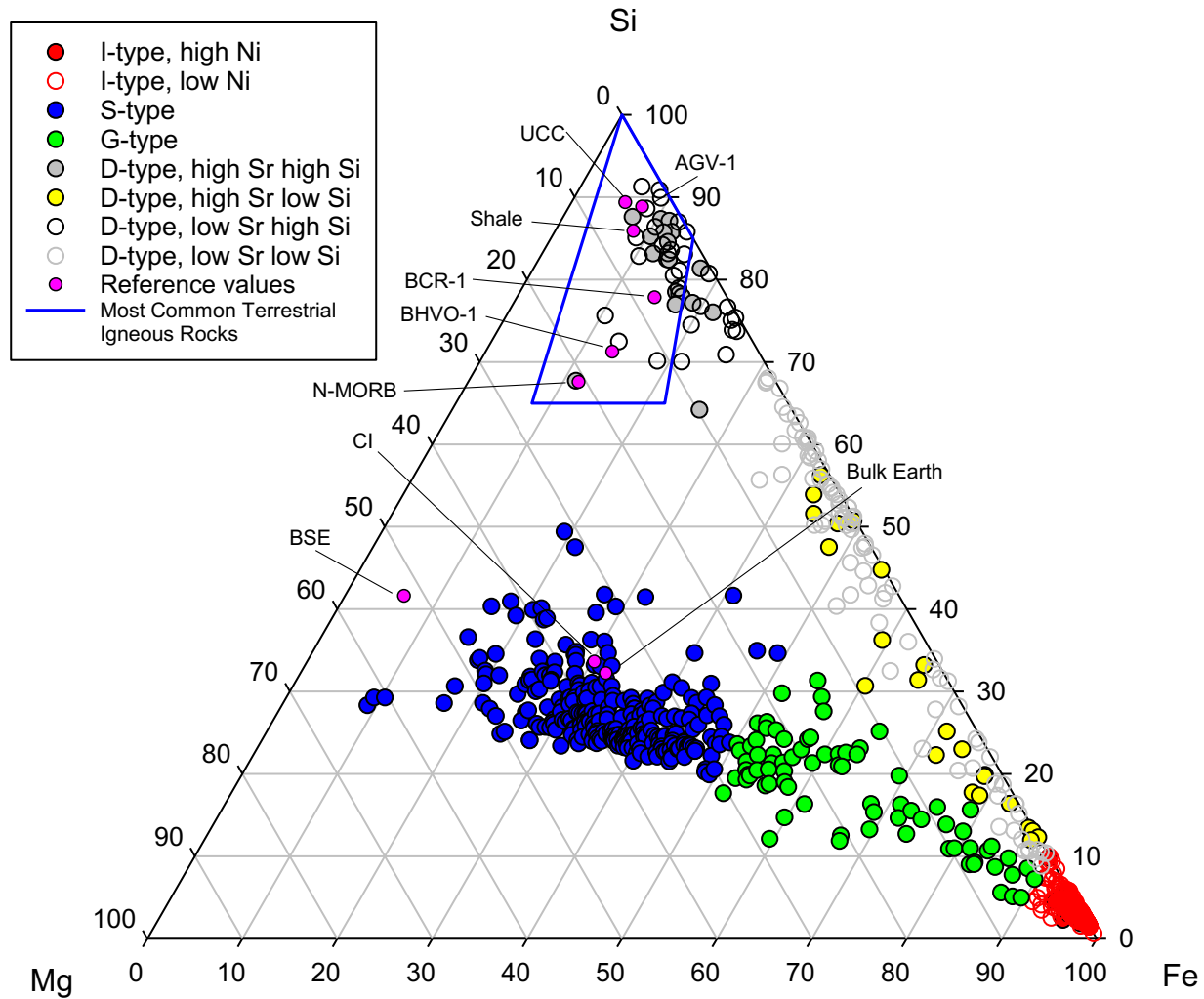


Figure 7. Atomic Mg-Si-Fe plot (atom %) of micro-XRF data for 745 IM1 site spherules and shards. The spherules are classified with the parameter values in **Table 1**. The spherule groups are compared to reference values of Earth materials (Bulk Earth, bulk silicate Earth [BSE], upper continental crust [UCC], shale, normal midocean ridge basalts [N-MORB], Hawaiian basalt [BHVO-1], Columbia River basalt [BCR-1], Guano Valley andesite [AGV-1]) and CI meteorites. Also shown is the range of chemical compositions of terrestrial igneous rocks.

For the primitive spherules we use $100\text{Fe}/(\text{Fe}+\text{Si}+\text{Mg}) > 90$ to distinguish I-types from S- and G-types. This group is supposed to be primarily made of iron compounds and is clearly distinct from the other primitive spherules in the $100\text{Fe}/(\text{Fe}+\text{Si}+\text{Mg})$ histogram (**Figure 8a**). This I-group is further subdivided into high Ni (>4000 ppm) and low Ni (<4000 ppm) groups (**Figure 8b**), as high Ni spherules are most likely of cosmic origin and typically have Ni > 4000 ppm (Engrand et al. 2006). The S- and G-type dividing line is $100\text{Si}/(\text{Fe}+\text{Si}+\text{Mg}) = 50$ (**Figure 8a**), relatively consistent with previous literature (cf. Brownlee et al. 1997; Taylor et al. 2000; Folco and Cordier 2015). The primitive spherule groups are compared to reference values of Bulk Earth, bulk silicate Earth (BSE) (McDonough and Sun 1995), and CI meteorites (Anders and Grevesse 1989) that are and should be on the primitive trend (**Figure 7**).

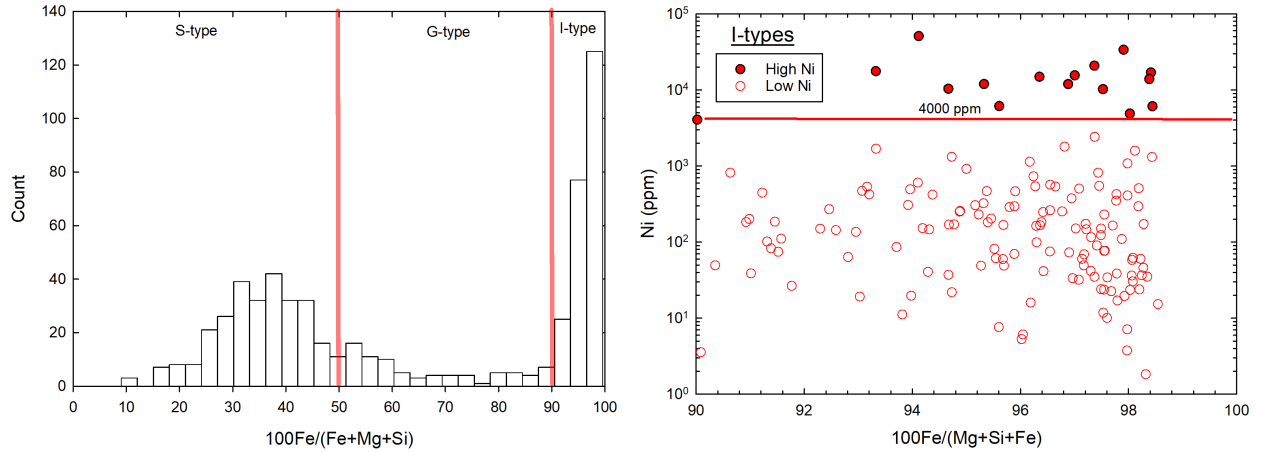


Figure 8. Left: Histogram of the $100\text{Fe}/(\text{Fe}+\text{Si}+\text{Mg})$ measured with micro-XRF for 745 IM1 site spherules. This diagram shows the dividing lines selected between S-, G-, and I-types. Right: Plot of Ni(ppm) vs. $100\text{Fe}/(\text{Fe}+\text{Si}+\text{Mg})$, showing the dividing line selected for high-Ni vs low-Ni spherules.

Owing to the incompatible nature of Sr, the Sr content of differentiated spherules is a good indication of the enrichment of refractory lithophile elements in the spherules and thus the extent of differentiation. The differentiated spherules in this study are thus divided into high Sr (>450 ppm) and low Sr groups (<450 ppm). They are further subdivided on the Si-content. For high Sr spherules we use $100\text{Si}/(\text{Fe}+\text{Si}+\text{Mg}) = 60$ as the dividing line and for low Sr spherules a value of $100\text{Si}/(\text{Fe}+\text{Si}+\text{Mg}) = 70$. These dividing lines are shown in **Figure 9**. The spherule groups are compared to reference values for samples of the Earth's crust in **Figure 7**. This includes average composition of normal midocean ridge basalts (N-MORB) (Gale et al. 2013), average upper continental crust (UCC) (Rudnick and Gao 2014) and shale (Ray and Paul 2021). Also plotted is the outline of the field of 37000 terrestrial igneous rocks as well as 3 USGS standards [Hawaiian basalt (BHVO-1), Columbia River basalt (BCR-1), Guano Valley andesite (AGV-1) (Jochum et al. 2007)].

We note that the high Si varieties of D-spherules plot close to or within the range of terrestrial igneous rocks, while the low Si groups do not. Thus, the D-type spherules have been divided into four distinct groups. This results in eight distinct spherule groups that are all shown in **Figure 7**. The parameters used to subdivide the spherules are summarized in **Table 1**. The total number of spherules identified in each group are listed in **Table 2**.

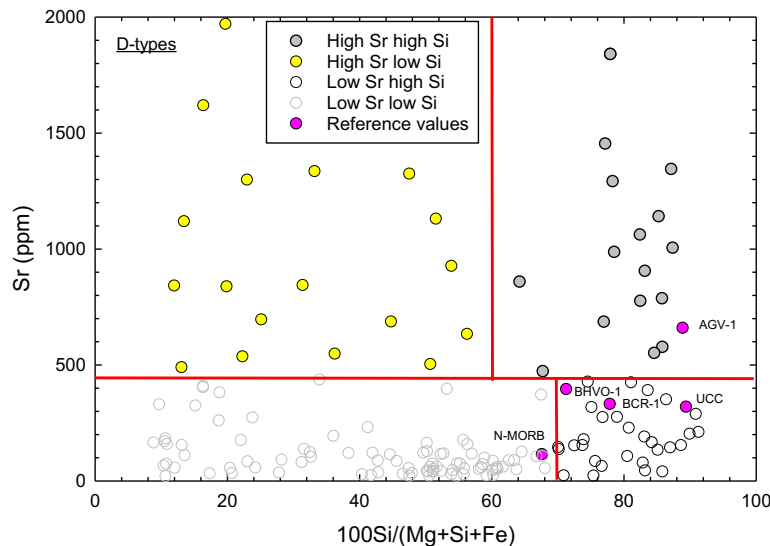


Figure 9. Micro-XRF data for 745 IM1 site spherules. Plot of Sr(ppm) vs. 100Si/(Fe+Si+Mg).

Table 1. Classification of the spherules with the Mg-Si-Fe diagram.

	Type	100Fe/ (Fe+Si+Mg)	Ni (ppm)	Mg/Si	Sr (ppm)	100Si/ (Fe+Si+Mg)
Primitive spherules	I-type, high Ni	>90	>4000			
	I-type, low Ni	>90	<4000			
	S-type	<50		>1/3		
	G-type	>50 and <90		>1/3		
Differentiated spherules	D-type, high Sr, high Si			<1/3	>450	>60
	D-type, high Sr, low Si			<1/3	>450	>60
	D-type, low Sr, high Si			<1/3	<450	>70
	D-type, low Sr, low Si			<1/3	<450	>70

Table 2. Comparison of the number of each spherule type identified by micro-XRF and by ICP-MS results.

	Type	Micro-XRF Number	Micro-XRF Percent	ICP-MS Number	ICP-MS Percent
Primitive spherules	I-type, high Ni	18	2.4	5	7.4
	I-type, low Ni	212	28.5	18	26.5
	S-type	275	36.9	19	27.9
	G-type	78	10.5	0	0
Differentiated spherules	D-type, high Sr, high Si	20	2.7	6	8.8
	D-type, high Sr, low Si	23	3.1	5	7.4
	D-type, low Sr, high Si	29	3.9	0	0
	D-type, low Sr, low Si	90	12.1	3	4.4
	D-type, BeLaU high Si			2	2.9
	D-type, BeLaU low Si			10	14.7
	Total	745	100	68	100

The major element and trace element compositions of 68 spherules from the IM1 site measured by ICP-MS are in **Table S2** (supplement). ICP-MS data for the spherules include many more elements than the micro-XRF data but does not include measurement of Si. Thus, we also plot the micro-XRF data in a Mg-Al-Fe diagram (**Figure 9a**) and compare it to the same diagram with ICP-MS measurements (**Figure 9b**). The micro-XRF results show the same groups in the Mg-Al-Fe plot as for the Mg-Si-Fe plot in **Figure 8**. Note that the primitive and differentiated trends are even better separated in **Figure 9a**, and the fields of the primitive spherule groups are also well defined by this diagram. We conclude that this new diagram is a good alternative to the Mg-Si-Fe diagram, when Si data are missing. The Sr concentrations of the D-type spherules measured by micro-XRF are plotted versus the Al/Fe ratio in **Figure 10a**. While there is some overlap, this diagram gives a relatively good separation of the high Si and low Si groups after separating the differentiated spherules into high (>450 ppm) and low (<450 ppm) Sr groups and is thus also useful for data with missing Si measurements. Classification of the spherules with the parameters in **Table 3** will thus yield a relatively equivalent classification of the spherules compared to that obtained with the Mg-Si-Fe diagram. The primitive spherules are identified by $100\text{Al}/(\text{Fe}+\text{Al}+\text{Mg}) < 10$ and the differentiated spherules by $100\text{Al}/(\text{Fe}+\text{Al}+\text{Mg}) > 10$. For the primitive spherules the I-types are identified by $100\text{Fe}/(\text{Fe}+\text{Al}+\text{Mg}) > 90$, the S-types by $100\text{Fe}/(\text{Fe}+\text{Al}+\text{Mg}) < 65$, and the G-types in between these two values.

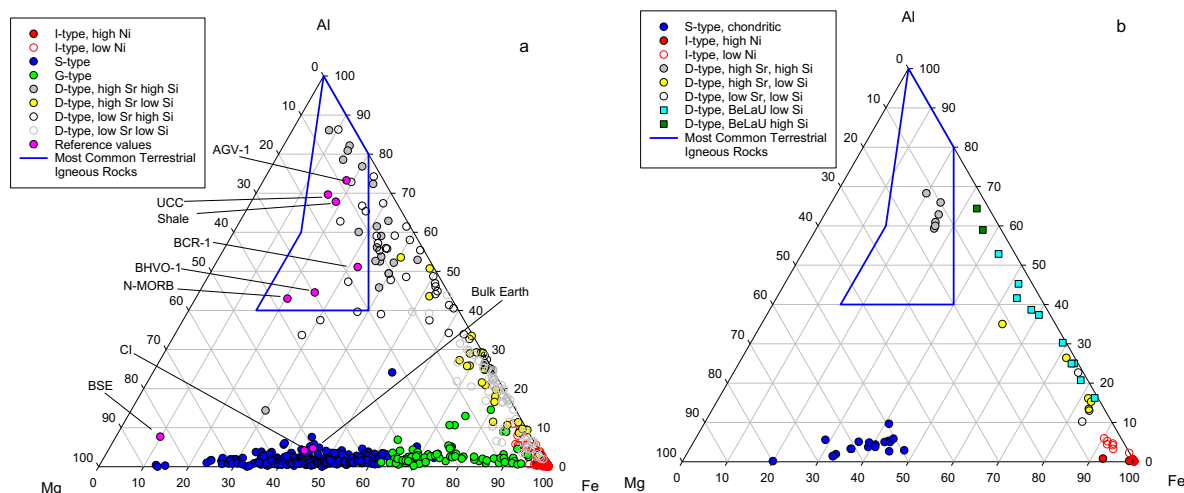


Figure 10. (Figure formatting, faint grey lines present as outlines) a) Atomic Mg-Al-Fe plot (atom %) of micro-XRF data for 745 IM1 site spherules. The spherules are classified with the parameter values in **Table 1** and are the same spherules shown in **Figure 7**. b) ICP-MS data for 68 spherules, classified with the parameters in **Table 3** as discussed in the text.

Table 3. Classification of the spherules with the Mg-Al-Fe diagram.

	Type	100Fe/ (Fe+Al+Mg)	Ni (ppm)	100Al/ (Fe+Al+Mg)	Sr (ppm)	Sr vs Al/Fe	BeLaU
Primitive spherules	I-type, high Ni	>90	>4000	<10			
	I-type, low Ni	>90	<4000	<10			
	S-type	<65		<10			
	G-type	>65 and <90		<10			
Differentiated spherules	D-type, high Sr, high Si			>10	>450	See Figure 3a	<80
	D-type, high Sr, low Si			>10	>450	See Figure 3a	<80
	D-type, low Sr, high Si			>10	<450	See Figure 3a	<80
	D-type, low Sr, low Si			>10	<450	See Figure 3a	<80
	D-type, BeLaU, high Si			>10			>80
	D-type, BeLaU, low Si			>10			>80

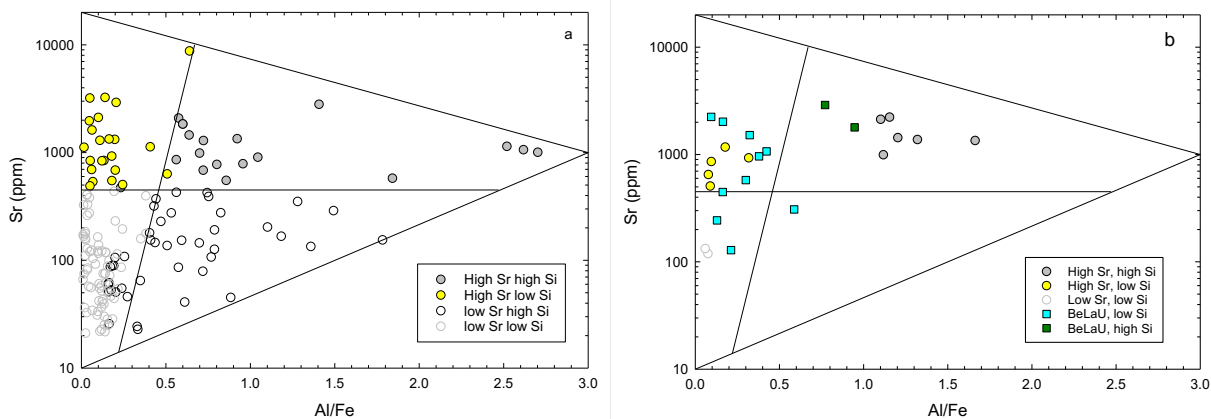


Figure 11. Sr (ppm) vs the Al/Fe weight ratio for D-type spherules from the IM1 site. a) Micro-XRF data, b) ICP-MS data.

For differentiated spherules we first use the fields of the four different D-types shown in **Figure 10a** to identify the same types for the ICP-MS data. In addition, we use **Figure 11** to identify spherules with particularly high contents of refractory lithophile elements, based on the enrichments of Be, La and U relative to Mg and Fe. First the concentrations of these elements are normalized to the same elements in CI chondrites. The CI-normalized values (Be_{CI} , La_{CI} , U_{CI} , Mg_{CI} and Fe_{CI}) are used to calculate the following plotting parameters for the ternary diagram in **Figure 12**:

$$BeLaU = \frac{Be_{CI} + La_{CI} + U_{CI}}{Be_{CI} + La_{CI} + U_{CI} + Mg_{CI} + Fe_{CI}}$$

$$M = \frac{1000(Mg_{CI})}{Be_{CI} + La_{CI} + U_{CI} + Mg_{CI} + Fe_{CI}}$$

$$F = \frac{100(Fe_{CI})}{Be_{CI} + La_{CI} + U_{CI} + Mg_{CI} + Fe_{CI}}$$

The BeLaU parameter was divided by 100, while the M-parameter was multiplied by 10 to make use of the entire area inside the ternary diagram. We define a BeLaU-type spherule as having $BeLaU > 80$, and we define high and low Si varieties based on the Al/Fe ratio, as in **Figure 11a**. This procedure identifies 10 of D-type spherules as low Si BeLaU-type spherules and 2 as high Si BeLaU-type spherules as shown in **Figure 12**. These spherules are also shown in **Figures 10b and 11b**.

BeLaU-type spherules can only be identified with ICP measurements. We found 12 BeLaU-type spherules out of 26 D-type spherules identified with ICP measurements (**Table 2**). The BeLaU spherules were found in tracks 4, 13, 14, 17 and 19. With the micro-XRF measurements, we identified 162 D-type spherules out of a total of 745 spherules. This results in an estimate of 10 % of all spherules to be of BeLaU-type. Thus, out of the ~802 spherule identified as natural materials by our micro-XRF and ICPMS measurements we estimate that there should be up to 80 BeLaU-type spherules in our collection.

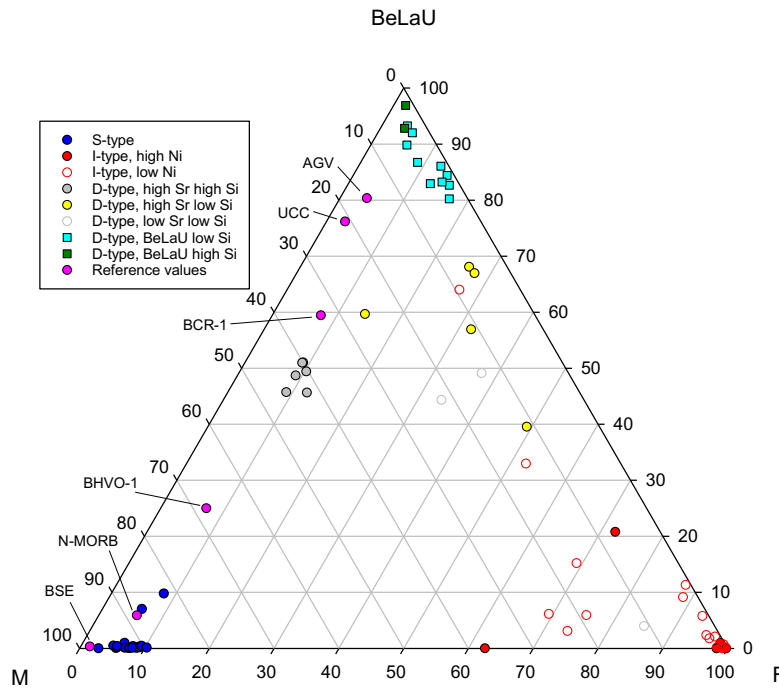


Figure 12. Ternary diagram for identifying the spherules that are most enriched in Be, La and U (BeLaU) relative to Mg (M) and Fe (F). The parameters are CI chondrite normalized (see text).

Origins of the spherule groups

In total we analyzed 850 magnetic particles by either micro-XRF (792) or ICP-MS (68) or both (11). We found that 48 have simple chemical compositions that likely make them derived from industrial products or terrestrial minerals. They were thus not included in this study and so we are left with 802 spherules for which we report data in this paper.

Origin of Primitive spherules: The primitive spherules are in general characterized by high Cr, Co, and Ni and highly variable Mn. Cr is moderately volatile, similar to Mn, thus in a Cr/Fe vs Mn/Fe plot (**Figure 13**) there is an overall positive correlation as expected for these ratios. We note that S-type spherules have both Cr/Fe and Mn/Fe similar carbonaceous chondrites (CCs) and non-carbonaceous chondrites (NCs), and there is no major effect from volatile loss of Cr and Mn for this type of spherules. The Ni-rich I-type spherules have very low Mn/Fe showing substantial volatile loss of Mn and also some for Cr. Most of the low Ni I-type spherules have normal chondritic Mn/Fe and slightly lower, so these do not have evidence for substantial volatile loss for Mn. This suggests that their low Cr content is due to some other partitioning process. The primitive nature of S- and I-type spherules are supported by their chondrite-like Cr/Fe and Mn/Fe, distinctly different from the planetary values (Earth, Moon, Vesta, Mars) shown in **Figure 13**. In a plot of Co/Fe vs Mn/Fe (**Figure 14a**) when comparing S-type spherules similar CCs and NCs, there is no major effect from volatile loss. When comparing with planetary bodies we see an overall negative correlation that is likely due to core formation in the planets. Ni-rich I-type spherules have very low Mn/Fe showing substantial volatile loss. Low Ni I-type again have normal chondritic Mn/Fe and slightly lower, so these do not have evidence for substantial volatile loss. The primitive nature of S- and I-type spherules are supported by their chondrite-like Co/Fe and Mn/Fe, distinctly different from planetary values. In a plot of Ni/Fe vs Mn/Fe (**Figure 14b**) we also see an overall negative correlation. Here the S-type spherules have lower Ni/Fe compared to CCs and NCs. Ni is more strongly partitioned into metal compared to Co, so some Ni rich metal must have been lost in producing the S-type Spherules. Thus, overall Mn, Cr, Co and Ni relative to Fe in S- and I-type spherules supports their origin from primitive chondritic meteorites but do show evidence of volatile loss as well as metal-silicate separation due the processing into spherules.

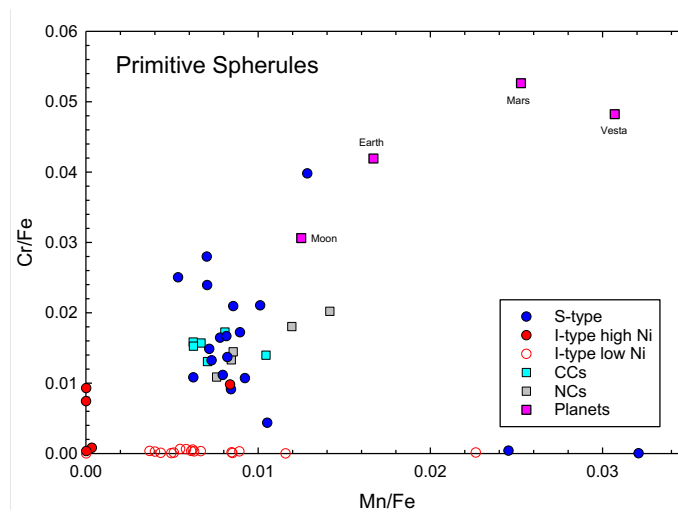


Figure 13. A plot of Cr/Fe vs Mn/Fe for primitive spherules compared to carbonaceous (CC) and non-carbonaceous (NC) chondrites, Earth, Mars, Vesta and the Moon. Sources: CCs and NCs from [Alexander](#)

2019ab. Mars: Yoshizaki and McDonough (2020), Vesta: Dreibus and Wänke (1980), Moon: Hauri et al. (2015).

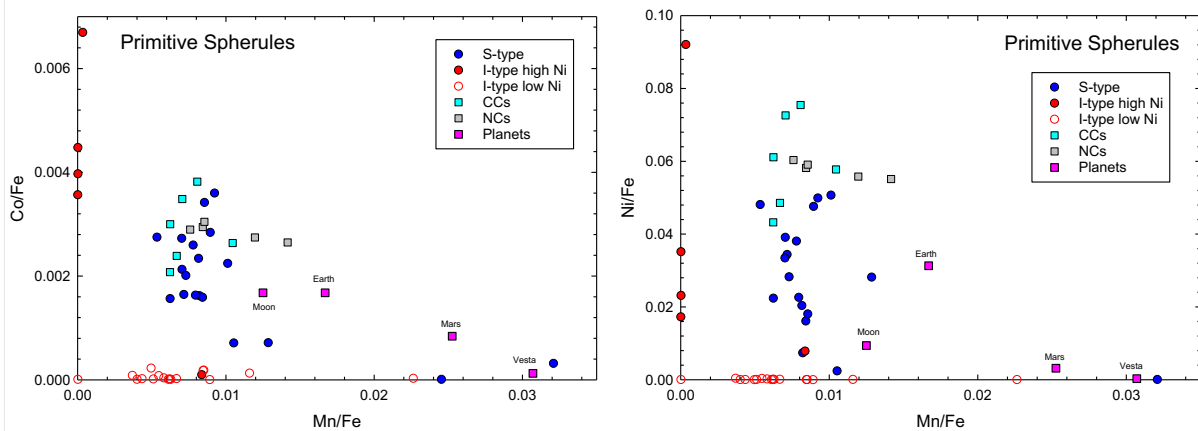


Figure 14. A plot of Co/Fe (a) and Ni/Fe (b) vs Mn/Fe for primitive spherules compared to carbonaceous (CC) and non-carbonaceous (NC) chondrites, Earth, Mars, Vesta and the Moon. Sources: CCs and NCs from Alexander 2019ab. Mars: Yoshizaki and McDonough (2020), Vesta: Dreibus and Wänke (1980), Moon: Hauri et al. (2015).

Origin of D-type spherules: The Mg-Al-Fe plot in **Figure 10** has also been used for classifying terrestrial komatiites, ultramafic lavas that are primarily found in the Archean (cf. Rickwood 1989). The only modification is that Fe is replaced with Fe+Ti. Such a plot is used in **Figure 15** to further understand the origin of D-type spherules. The importance of this diagram is that it shows komatiites as separate fields from tholeiitic basalts and calc-alkaline and tholeiitic magma series. The blue outline shows the field of common terrestrial rock while the red outline includes all terrestrial igneous rocks. The importance for interpreting the D-type spherules is that all except some of the Si-rich D-type spherules are outside the field of common terrestrial igneous rocks and most of them are outside the red outline where there are no terrestrial igneous rocks. In particular, the low-Si BeLaU spherules are all close to the expected meteor path, except for two outside this field. The BeLaU spherules are also clearly different from the lunar high-K KREEP composition, which is the magmatic composition on the Moon with the highest concentration of Be, La and U. Martian SNC meteorites reflect the Fe-rich composition of the martian mantle and are the only known solar system igneous rocks that plot in the triangle between the red line and the Fe+Ti apex of the diagram. However, there are no Martian igneous rock that get as close to the Al-(Fe+Ti) side of the diagram as the BeLaU-type spherules. The BeLaU-type spherules composition could probably be produced by extreme differentiation of an iron-rich Martian-like magma that produced the igneous rocks among the Martian SNC meteorite.

BeLaU spherules are observed to be depleted in volatile elements such as Zn, Mn, K, and Pb. In addition, potassium and Mn are both volatile during planet formation while U and Fe are more refractory, so planets end up with different K/U and Mn/Fe ratios. The data for the D-type spherules are shown in a K/U vs Mn/Fe diagram (**Figure 16**). The reason for using these elemental ratios is that they both have for a long time been used to distinguish materials from different bodies in the solar system (cf. Papike et al. 2017; Halliday and Porcelli 2001). This works because K and U are both highly enriched in melts, so their ratio does not change much during igneous processes. Therefore, unaltered igneous rocks typically directly reflect a ratio close to the planetary value. In contrast Fe and Mn are both roughly equally distributed between melts and solids, so their ratio also does not change during igneous processes. Since K and Mn are both volatile one would expect both to be lost from the spherules during their flight through the atmosphere. The D-type spherules are compared with primitive chondritic meteorite and planetary values in **Figure 16**. There are three distinct groups. The group A spherules are similar to the Earth values and could by this comparison represent some terrestrial volcanic material. We note that these are all Si-rich spherules which

also make them similar in composition to terrestrial igneous rocks. Group B is strongly depleted in K relative to the Earth. Group C is strongly depleted in both K and Mn relative to the Earth. Most of the BeLaU spherules are in the group C field and none are in the Earth-like group A field. This shows that all the BeLaU spherules have exceptionally high volatile element loss, suggestive of extraterrestrial origin.

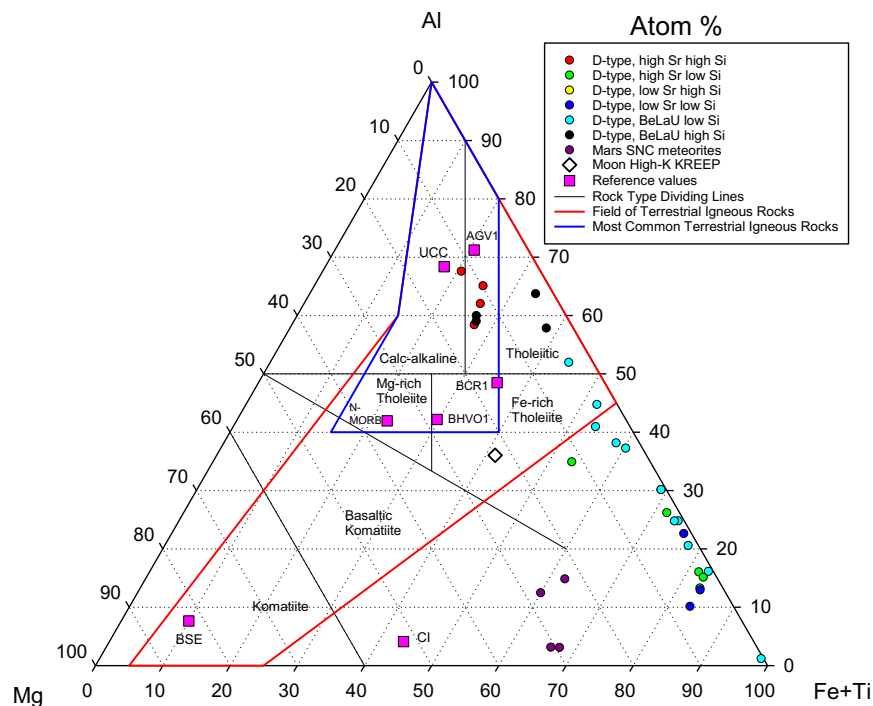


Figure 15. Atomic Mg-Al-Fe plot (atom %) Comparison of IM1 site spherules to differentiated materials from Earth (same references as for **Figure 7**), Moon (High-K KREEP; [Warren \(1989\)](#)) and Mars (SNC meteorites; [Lodders \(1998\)](#)).

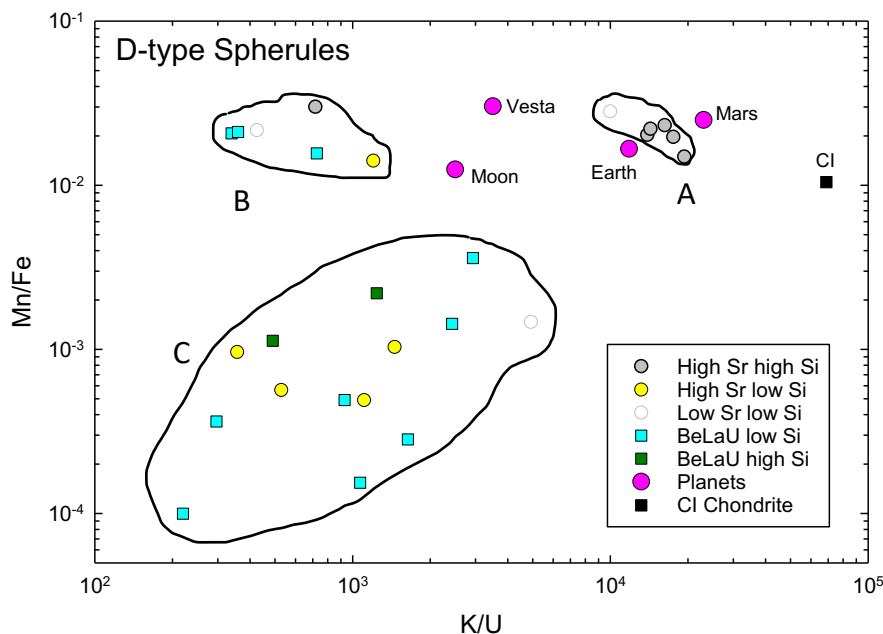


Figure 16. A plot of Mn/Fe vs K/U for D-type spherules compared to CI chondrites, Earth, Mars, Vesta and the Moon. Sources: [Anders and Grevesse 1989](#); [McDonough and Sun 1995](#) Mars: [Yoshizaki and McDonough \(2020\)](#), Vesta: [Dreibus and Wänke \(1980\)](#), Moon: [Hauri et al. \(2015\)](#).

Origin of the novel BeLaU-type spherules: The spherules with enrichment of beryllium (Be), lanthanum (La) and uranium (U), labeled “BeLaU”, appear to have an exotic composition different from other Solar System materials. The results for the “BeLaU” spherule S21 are displayed in **Figures 17a and 17b**. We use this spherule to point out some of the unique feature of the “BeLaU” elemental composition. A plot of the elemental abundances of the spherule S21 (normalized to CI chondrites) as a function of atomic number for 56 elements is shown in **Figure 17a**. Across the diagram the peak abundances are for Be, La and U, hence the name “BeLaU.” The abundance pattern of S21 implies derivation from a planetary crust, highly enriched in refractory lithophile elements (red dots). The volatile element abundances (green dots) are very low, suggesting either derivation from an extremely volatile-depleted planet or evaporative loss during passage through the Earth’s atmosphere. The very low content of refractory siderophile elements with affinity to iron (Re) suggest a source planet with an Fe core. Since there are strong indications that the spherules are derived from a differentiated planet, the data are also plotted in **Figure 17b** as a function of an igneous compatibility sequence. Compatibility is a geochemical parameter measuring how readily a particular element substitutes for a major element in mantle source minerals during melting to produce magma. It also roughly represents the sequence of enrichments of elements in a crystallizing magma.

The abundances of elements as function of their compatibility for all 12 “BeLaU”-type spherules are shown in **Figure 18**. The “BeLaU” spherules’ variations in the abundances of trace elements relative to CI chondrites are higher by 1-3 orders of magnitude compared to cosmic spherules from the solar system reviewed by [Folco et al. \(2015\)](#). The “BeLaU” element patterns are compared in **Figure 19** with the highly enriched sources of differentiated bodies in the solar system, including Earth’s upper continental crust ([Rudnick and Gao, 2014](#)) and kimberlites ([Giuliani et al., 2020](#)), lunar magma ocean residual liquid (KREEP) ([Warren, 1989](#)), Shergottites from Mars ([Lodders, 1998](#); [Jambon et al., 2002](#)) and eucrites from Vesta ([Kitts and Lodders, 1998](#)). **Figure 19** presents the comparison of refractory lithophiles that are not easily lost by evaporation or altered by fluids. Shergottites and eucrites are markedly distinct from the “BeLaU” samples due to the systematically lower incompatible element enrichments. The upper continental crust is also overall depleted in incompatible elements compared to “BeLaU.” Additionally, “BeLaU” samples have prominent negative anomalies of Ti, Li, higher Lu/Al, and variable Be and Sr enrichments that do not match the smooth pattern of the upper continental crust. Kimberlite is also remarkably distinguished from the “BeLaU” pattern in their Ta and Nb positive anomalies and the strong heavy rare earth element depletion. Lastly, despite the resembling overall element enrichments, the lunar KREEP displays pronounced differences in light rare earth element enrichment and strong Sr, Eu, and Cr negative anomalies from “BeLaU” that distinguish the two groups.

Figure 20 compares the BeLaU elemental pattern with the average upper continental crust composition. This shows that BeLaU-type spherules are depleted by a factor of 2 to more than 10 for all volatile elements (Na, K, Mn, Zn, Rb, Cs, Tl, Pb and Bi), while refractory lithophile elements are enriched by a factor of 3-20 compared to the average continental crust. Also, the rare earth element pattern is substantially fractionated when compared to the average upper continental crust. We must conclude that the elemental pattern of the BeLaU spherules is very different from the upper continental crust composition. The volatile elements (such as K, Mn, Zn and Pb) were most likely lost by evaporation during the passage of a bolide like IM1 through the Earth’s lower atmosphere.

We pointed out earlier the difference of the BeLaU signature from KREEP in major elements (**Figure 15**). In spite of some distinct differences from KREEP there are enough similarities in the elemental composition to consider a similar type of origin for both patterns. The lunar KREEP composition is widely believed to be the result of crystallization of an early lunar magma ocean ([Snyder et al., 1992](#); [Rapp and Draper, 2018](#); [Charlier et al., 2018](#)). We conclude that the “BeLaU” samples possibly reflect a highly differentiated and extremely evolved composition, but from an unknown source.

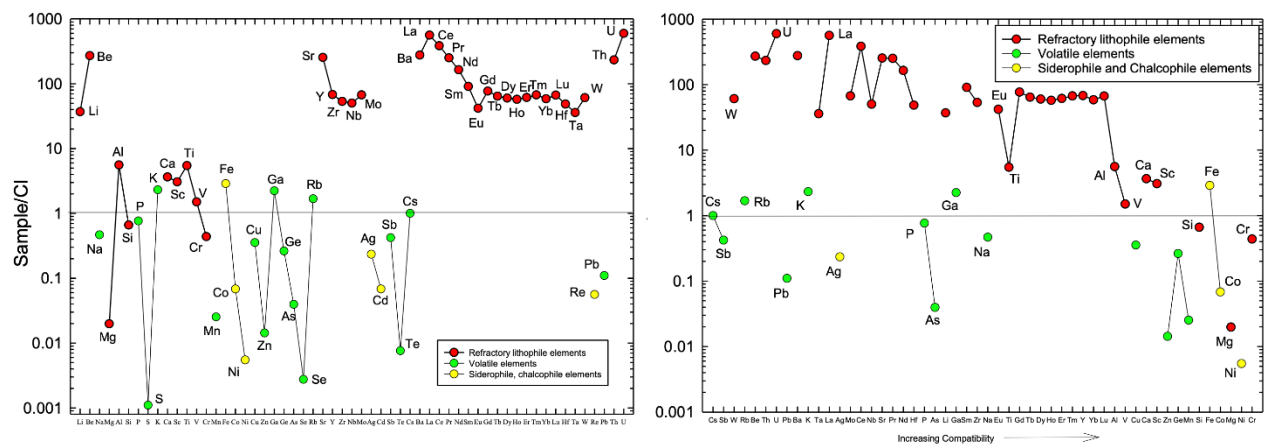


Figure 17: a) The “BeLaU” elemental abundances of the spherule S21 (normalized to CI chondrites) versus atomic number for 56 elements. The solar system standard of CI chondrites is represented by a value of unity on the plot. b) Abundances versus compatibility (see text) for S21. Elements ordered with increasing compatibility towards the right.

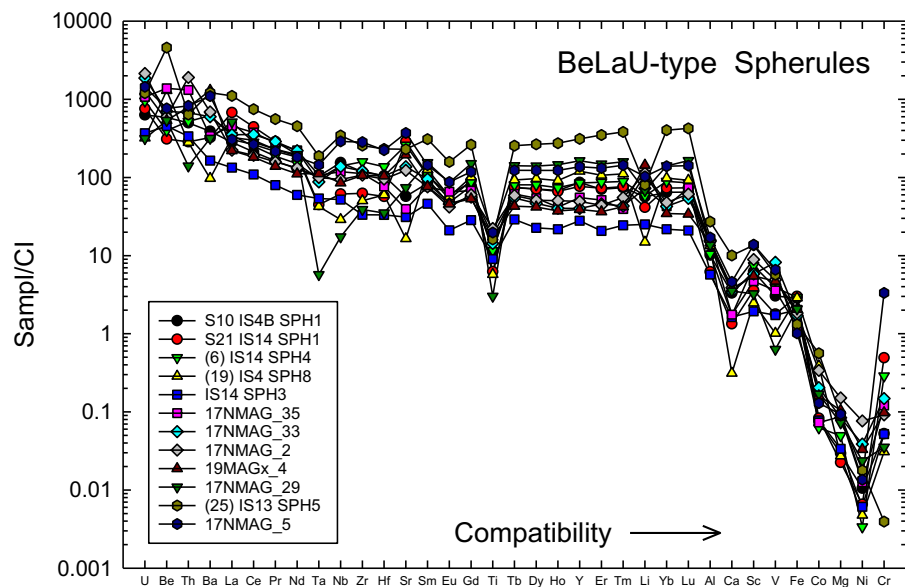


Figure 18. Abundances of refractory lithophile elements as function of their compatibility for all 12 “BeLaU”-type spherules.

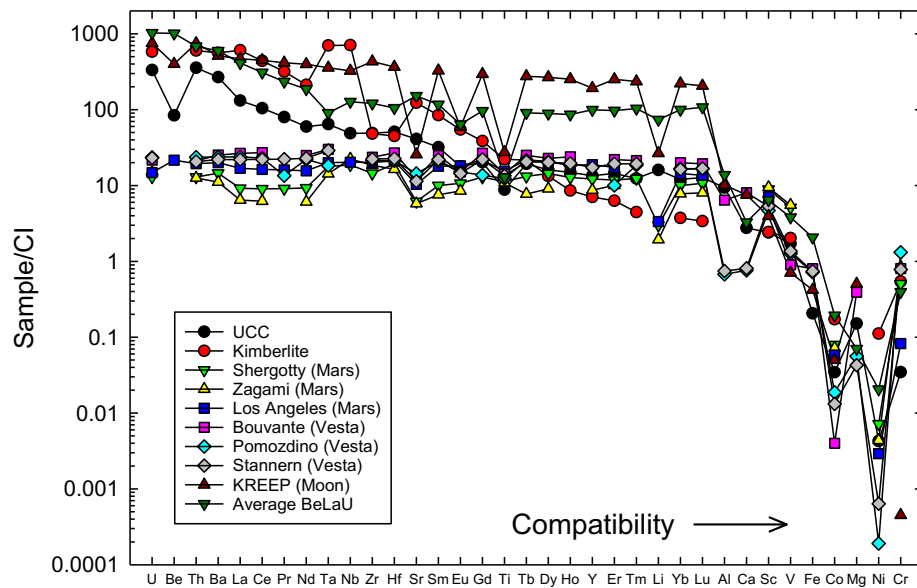


Figure 19. The “BeLaU” element abundance and pattern distinguish from the enriched sources from the solar system bodies. Elements are ordered with increasing compatibility. Data source: Earth’s upper continental crust ([Rudnick and Gao, 2014](#)); kimberlites ([Giuliani et al., 2020](#)); lunar KREEP ([Warren, 1989](#)); shergottites (Shergotty, Zagami, and Los Angeles) ([Lodders, 1998](#); [Jambon et al, 2002](#)); eucrites (Bouvante, Pomozdino, and Stannern) ([Kitts and Lodders, 1998](#)).

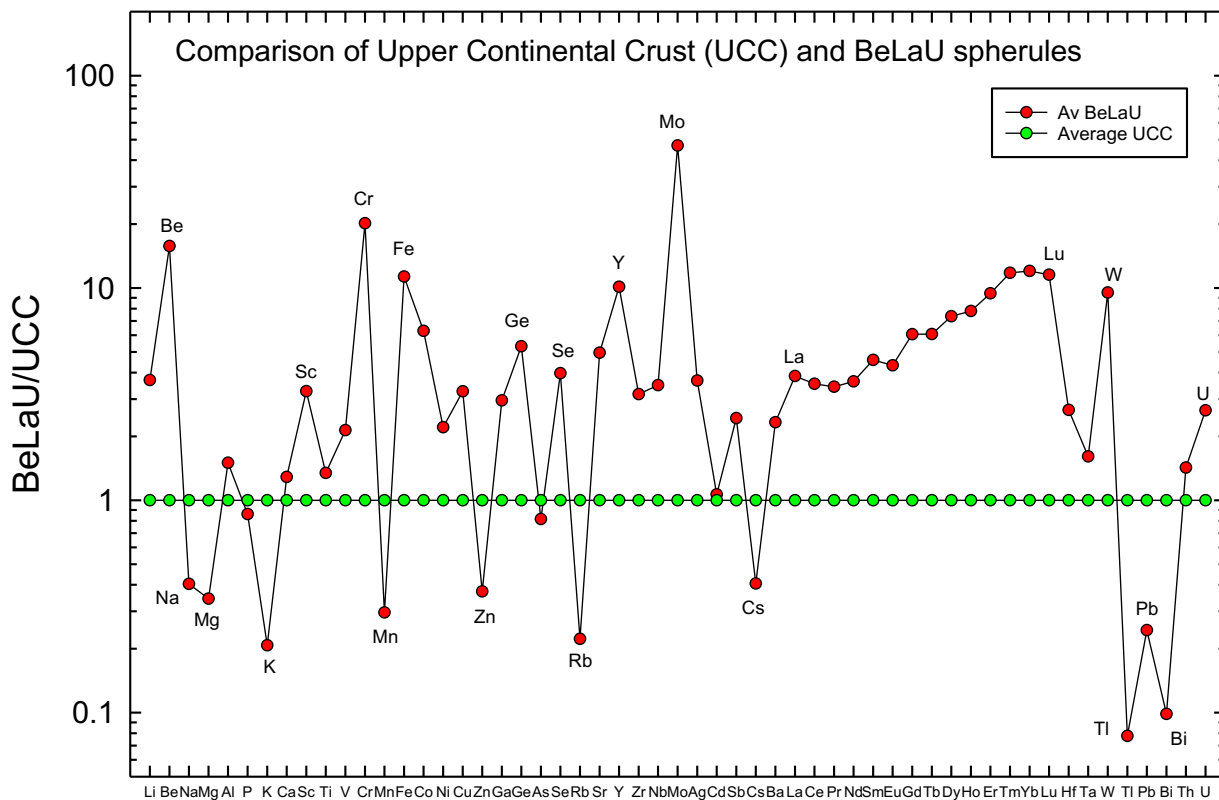


Figure 20. Comparison of average BeLaU-type spherules with average upper continental crust.

Conclusions

The magnetic sled survey around IM1's path during 14–28 June 2023 resulted in 850 magnet particles (consisting mostly of spherules) of diameters within the range of 0.1–1.3 millimeters through 26 runs surveying 0.06 km². It is not yet clear as to whether any of these specimens are associated with the IM1 bolide. We identified four distinct groups of primitive spherules (S-type, G-type, I-type high Ni, I-type low-Ni). We also identified a substantial fraction (22%) of the spherules as being derived from igneous precursors. As these are clearly not related to typical achondrite meteorites as for previously discussed differentiated spherules we denote these new types of differentiated spherules as D-type spherules. The D-type spherules were classified into four distinct groups. The chemical compositions of 10 % of the spherules show extremely strong enrichment of refractory lithophile elements such as Be, La and U (BeLaU-type spherule), but very low refractory siderophile elements such as Re. Volatile elements, such as K, Mn, Zn and Pb are very low in the BeLaU-type spherules and were most likely lost by evaporation during passage through the Earth's lower atmosphere. While the BeLaU-type spherules clearly appear to be derived from material formed by igneous fractionation, their chemical composition is novel in that there it is different from known existing solar system materials that have been analyzed thus far, with the KREEP component of the lunar crust being closest. The “BeLaU” samples reflect a highly differentiated, extremely evolved composition, of an unknown origin.

These interpretations will be considered critically along with additional results from spherule analysis in future publications. Recovering larger pieces of similar composition from the search area and comparing the findings to distant control regions would certainly help in the interpretation.

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Competing interests. No competing interests.

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Supplementary Material.

Data tables in Excel spreadsheets:

Table S1. Major and trace elements concentrations (ppm) measured by micro-XRF for 745 samples (data plotted in Figures 5, 6, 7, 8, 9a, 10a).

Table S2. Major elements, Be, Cr, Co, Ni, Sr, La and U concentrations (ppm) measured by ICP-MS for 68 samples (data plotted in Figures 9b, 10b, 11, 12, 13, 14, 15).

Table S3. Elemental data (ppm) measured by ICP-MS for 12 BeLaU-type spherules and the CI normalizing values of Anders and Grevesse (1989).

Table S4. Electron microprobe data on the chemical composition in the different region on the surface of the spherule S21, as labeled in **Figure S1**.

Element	Pt#1 (weight %)	Pt#2 (weight %)	Pt#3 (weight %)	Pt#4 (weight %)	Pt#5 (weight %)	Pt#6 (weight %)	Pt#7 (weight %)
C	0.64		0.7	1.5	2.1	5.48	2.66
O	13.95	28.9		32.35	29.18	18.66	34.19
Na	0.63	0.73			0.7	0.72	
Mg	0.28				0.42	0.15	
Al	2.93	3.48		1.15	2.45	1.44	1.18
Si	6.4	10.85			2.92	4.2	0.97
P		0.31					
Cl	0.26	0.4			0.51	1.28	
K	0.39						
Ca	3.88	4.55			1.73	1.85	0.23
Cr					0.31		
Fe	70.65	50.78	99.3	65.01	59.69	66.22	60.77
Total	100	100	100	100	100	100	100

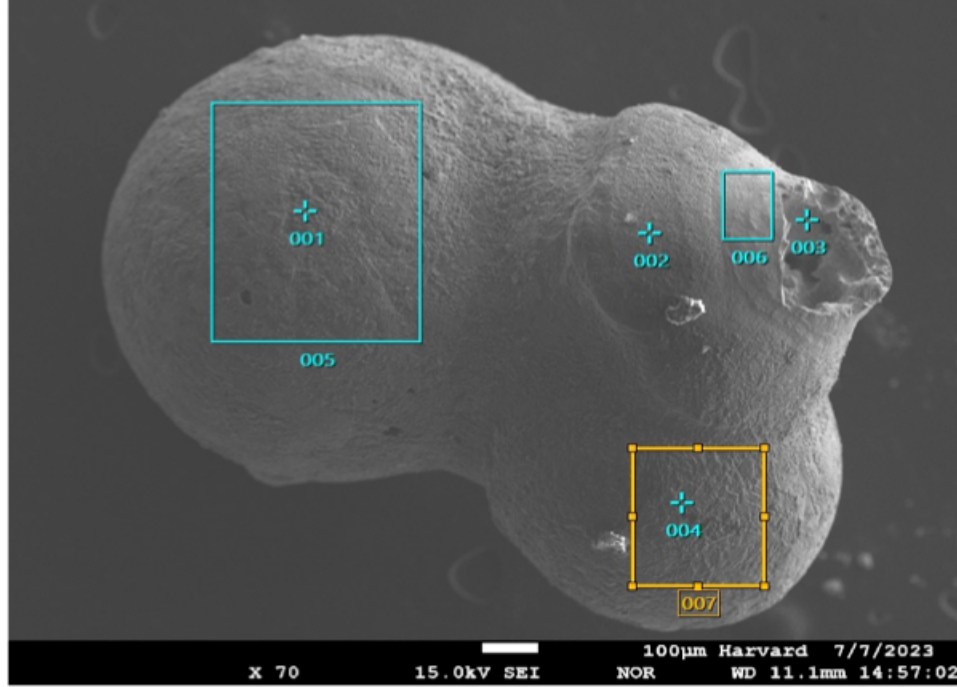


Figure S1. An example of a large (1.3 mm in maximum diameter, 0.9 mm average) spherule in the region along IM1's path is S21 (IS14-SPH1) from run 14.

An example of a large (1.3 mm in maximum diameter, 0.9 mm average) spherule in the region along IM1's likely path is S21 (IS14-SPH1) from run 14. This lopsided spherule, shown in **Fig. S1**, is a composite of three spherules that solidified shortly after merger but too late for the merger product to become spherical. The mass of S21 (1.7 mg) is about twice that of IS16A SPH1 (0.84 mg). The existence of a triple-merger like S21 can potentially be explained as a product of a meteoric airburst. The total mass collected by S21-like spherules in all our runs is of order ~ 0.1 g. Given the sled's width of 1 m, the total surveyed area, ~ 0.06 km², constitutes a fraction of $\sim 10^{-3}$ of IM1's strewn field. This implies a total mass in S21-like spherules in our estimated strewn field of order ~ 100 g, as expected given that most of IM1's mass evaporated to undetectable particles (well below tens of μ m in size) or gas (Tillinghast-Raby et al., 2022). Assigning a mass of $\sim 10^{-3}$ g per spherule implies a total number of $\sim 10^5$ such spherules. Based on IM1's speed and fireball energy, the total mass ablated by IM1's fireball is $\sim 5 \times 10^5$ g corresponding to an object radius of $R \sim 50$ cm (Siraj and Loeb, 2022b). The total number of spherules divided by the initial volume associated with R yields an initial spherule number density of $n \sim 0.2$ cm⁻³ which gets diluted as the material expands. For the characteristic diameter of a spherule, ~ 1 mm, the geometric cross-section for spherule-spherule collisions is $\sigma \sim 4 \times 10^{-2}$ cm². The resulting collision probability is $\tau \sim n\sigma(2R) \sim (0.2 \text{ cm}^{-3}) \times (4 \times 10^{-2} \text{ cm}^2) \times (100 \text{ cm}) \sim 0.8$, implying a likelihood of $\tau^2 \sim 0.6$ for triple-spherule mergers such as S21. Mergers that occur inside a liquid envelope would result in a spherical shell with embedded sub-spherules inside of it.

Additional images of spherules are shown in Figure S2.

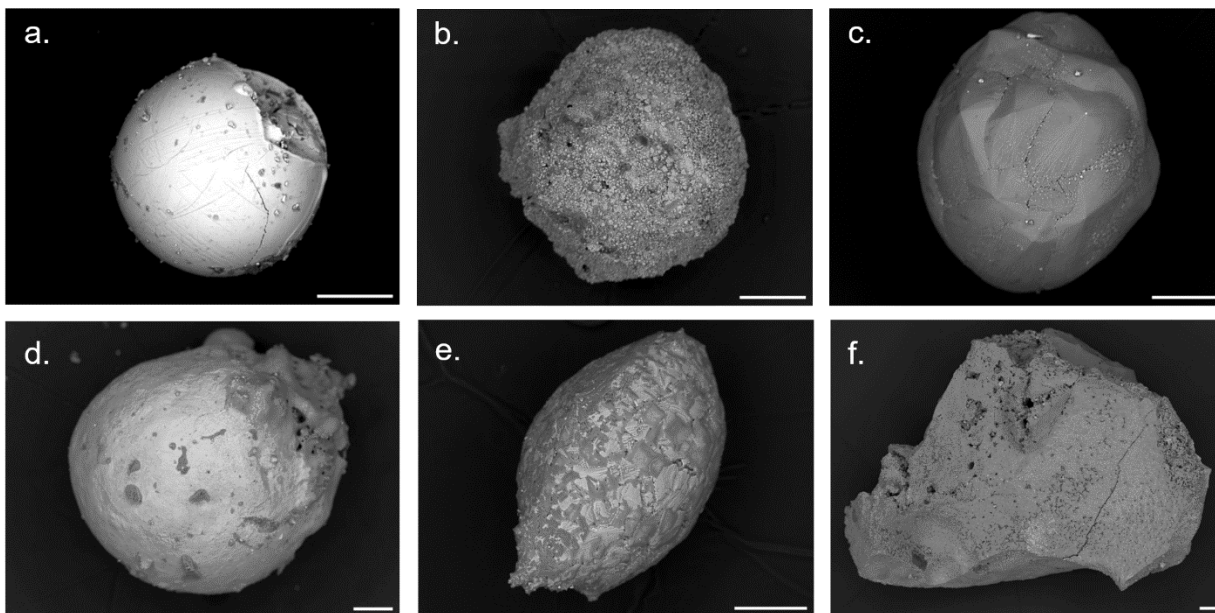


Figure S2. BSE images displaying varied morphologies of the samples in this study: a) IS8M2-1 (I-type, low Ni), b) IS13M1-19 (D-type, low Sr, low Si), c) IS20M-18 (S-type spherule displaying a polyhedral or “turtle-back” morphology), d) IS22M-16 (D-type, high Sr, low Si), e) IS14M(A)-7 (S-type spherule), and f) IS22M1-18 (identified as “shard,” no compositional classification). The scale bars indicate 100 microns.

Comparison to Coal Fly Ash and Spherule Data Reported by Rudraswami et al. (2016)

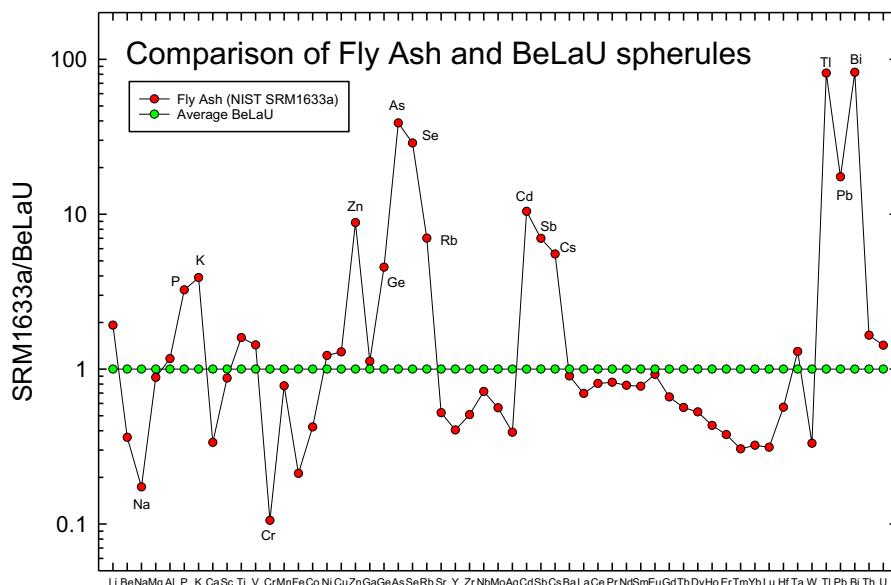


Figure S3. Comparison BeLaU with the NIST coal fly ash standard SRM1633a for 55 elements.

It has been claimed that the compositions of BeLaU spherules are consistent with coal ash ([Gallardo 2023](#)). The National Institute of Standards and Technology (NIST) has provided standards of coal fly ash. The best documented standard for many elements is SRM 1633a ([Jochum et al. 2005](#)). We compare the

average composition of BeLaU spherules for 55 elements with the SRM1633a coal fly ash standard in **Figure S3**. Many volatile elements (Zn, As, Se, Cd, Tl, Pb and Bi) are enriched in the coal fly ash by factors of about 10 to 100 compared to the BeLaU spherules. Some refractory elements (Be, Ca, Cr, Fe, Y, Tm, Yb, Lu W) are depleted by factors of 3 to 10 in coal fly ash when compared to BeLaU spherules. Thus, BeLaU spherules do not have the composition of coal fly ash, making the claim of [Gallardo \(2023\)](#) invalid ([Loeb et al. 2024a](#)). The BeLaU composition shows an excess of Be, La and U, and other elements by up to three orders of magnitude relative to the solar system standard of CI chondrites.

Many of these elements are 1–2 orders of magnitude more abundant than for the spherules reported by [Rudraswami et al. \(2016\)](#) (their **Figure 2**). All spherules, except for one, reported by [Rudraswami et al. \(2016\)](#) are S-type spherules (**Figure S4**), and the one exception does not have trace element enrichments like our BeLaU spherules.

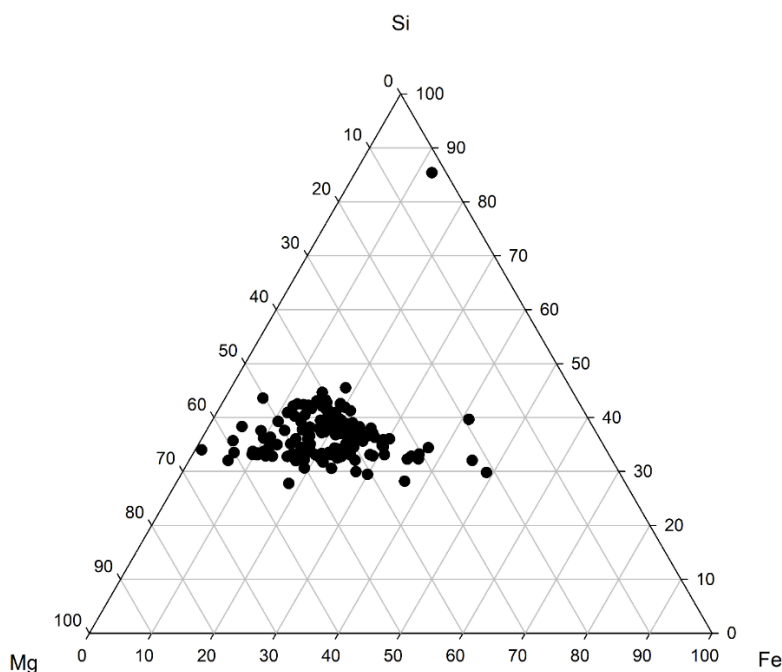


Figure S4 Ternary Mg-Si-Fe plot (atom %) of spherules from [Rudraswami et al. \(2016\)](#) normalized to (Mg + Fe + Si) and plotted in elemental proportions. The spherules in this study have been characterized to be S-type spherules.

Table S1. Major and trace elements concentrations (ppm) measured by micro-XRF for 745 samples (data plot

Type	Sub-type	Track	Sample name	O	Na	Mg
I-type	High Ni	15	IS 15 MAG-1__5_2	274729	15459	340
I-type	High Ni	6	IS 6_1	304404	0	1663
I-type	High Ni	19	19MAGx_25	291485	0	576
I-type	High Ni	8	IS 8-MAG-2__20_1	297798	1999	2288
I-type	High Ni	12	IS 12-MAG-2__24_2	303632	15843	6400
I-type	High Ni	12	IS 12-MAG-2__25_1	303632	15843	6400
I-type	High Ni	12	IS 12 -MAG 1__32_1	297950	0	43
I-type	High Ni	13	IS 13M UNSIFTED 1 s__14_1	301568	0	0
I-type	High Ni	13	IS 13M UNSIFTED 60 s__14_1	302748	3873	1165
I-type	High Ni	12	IS 12 -MAG 1__15_1	298838	0	374
I-type	High Ni	14	IS 14M Unsifted__22_1	301799	0	7305
I-type	High Ni	22	IS 22-MAG -1__12_1	302124	3920	900
I-type	High Ni	12	IS 12-MAG-2__8_1	305536	9908	1127
I-type	High Ni	14	IS 14M Unsifted__42_1	302032	2072	195
I-type	High Ni	22	IS 22-MAG-3__8_1	300689	15033	4190
I-type	High Ni	24	IS 24 -MAG__4_1	300929	4216	0
I-type	High Ni	12	IS 12 -MAG 1__10_1	302259	0	201
I-type	High Ni	12	IS 12-MAG-2__10_1	334652	23474	475
I-type	Low Ni	17	17MAG_37	300569	1247	1270
I-type	Low Ni	15	IS 15 MAG-1__3_1	304121	23761	0
I-type	Low Ni	10	IS 10 -MAG__3_1	307534	8489	8556
I-type	Low Ni	13	IS 13-M-1__11_1	303294	689	38
I-type	Low Ni	16	IS 16 MAG_2__13_2	306420	3088	3762
I-type	Low Ni	11	IS 11-MAG__13_1	303069	0	0
I-type	Low Ni	12	IS 12 -MAG 1__33_1	306796	1290	1315
I-type	Low Ni	20	IS 20 MAG__22_1	303053	721	508
I-type	Low Ni	12	IS 12-MAG-2__26_1	307734	9472	0
I-type	Low Ni	14	IS 14M Unsifted__38_1	304500	2098	608
I-type	Low Ni	20	IS 20 MAG__33_2	317672	8769	107
I-type	Low Ni	16	IS 16 MAG_2__2_2	305332	11039	589
I-type	Low Ni	17	17MAG_42	306544	2546	3681
I-type	Low Ni	19	19NMAG_3	304372	753	1487
I-type	Low Ni	24	IS 24 -MAG__14_1	303837	1307	1283
I-type	Low Ni	19	IS 19 M__23_1	305791	7426	359
I-type	Low Ni	19	IS 19 M__24_1	301555	17591	3086
I-type	Low Ni	12	IS 12-MAG-2__23_1	316414	13455	1868
I-type	Low Ni	12	IS 12 -MAG 1__9_2	303406	0	551
I-type	Low Ni	24	IS 24 -MAG__1_1	304920	660	1258
I-type	Low Ni	16	IS 16 MAG_2__17_1	313828	2688	5057
I-type	Low Ni	14	IS 14M UNSIFTED B 60 s_5_1	309689	7552	7565
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__27_1	306750	1912	2991
I-type	Low Ni	19	19NMAG_46	305305	0	1566
I-type	Low Ni	15	IS 15 MAG-1__4_1	317903	8047	3985
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__40_1	305544	709	319
I-type	Low Ni	20	IS 20 MAG__32_1	310137	3592	3198

I-type	Low Ni	13	IS 13M UNSIFTED 1 s__37_1	316618	629	6424
I-type	Low Ni	8	IS 8-MAG-2__6_2	301860	4121	109
I-type	Low Ni	24	IS 24 -MAG__7_1	306571	3804	0
I-type	Low Ni	12	IS 12-MAG-2__2_2	303385	6386	44
I-type	Low Ni	17	IS 17-MAG__4_1	308712	522	1698
I-type	Low Ni	17	17NMAG_30	300436	25	1427
I-type	Low Ni	22	IS 22-MAG-3__9_2	308601	1625	649
I-type	Low Ni	19	IS 19 M__8_1	301286	598	0
I-type	Low Ni	19	IS 19 M 24 Small	302242	22600	3702
I-type	Low Ni	12	IS 12 -MAG 1__5_1	306791	2481	1814
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__37_1	317105	5035	11639
I-type	Low Ni	13	IS 13-M-1__1_1	302524	12630	203
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__12_1	311555	3571	6895
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__24_1	304994	2598	600
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__27_1	306739	12459	3969
I-type	Low Ni	14	IS 14M Unsifted__15_1	304205	10353	734
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__12_1	312012	0	6130
I-type	Low Ni	9	IS 9C S18	303107	413	343
I-type	Low Ni	20	IS 20 MAG__31_1	302571	71764	1066
I-type	Low Ni	19	IS 19 M__15_1	316972	44498	12078
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__25_1	325212	83	6902
I-type	Low Ni	8	IS 8-MAG-2__29_2	317850	83137	4147
I-type	Low Ni	8	IS 8-MAG-2__14_1	304988	1289	1685
I-type	Low Ni	19	IS 19 M__11_1	308426	1331	1121
I-type	Low Ni	12	IS 12 -MAG 1__1_1	305523	0	953
I-type	Low Ni	12	IS 12-MAG-3__8_1	302331	3397	0
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__61_1	308334	5484	2485
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__54_1	306127	16085	3746
I-type	Low Ni	19	19MAG_3	305049	0	2051
I-type	Low Ni	17	17MAG_44	304852	1048	1630
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__40_1	305830	0	0
I-type	Low Ni	15	IS 15 M 2__1_1	307071	26682	192
I-type	Low Ni	20	IS 20 MAG__19_2	308853	3007	4221
I-type	Low Ni	12	IS 12 -MAG 1__39_2	303983	0	2389
I-type	Low Ni	15	IS 15 MAG-1__7_2	300615	68362	0
I-type	Low Ni	10	IS 10 -MAG__1_1	310690	25911	2354
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__56_1	304021	3135	488
I-type	Low Ni	22	IS 22 MAG-2__2_1	311239	6443	1115
I-type	Low Ni	19	IS 19 M__10_1	313124	26606	3863
I-type	Low Ni	12	IS 12-MAG-3__3_1	311601	30776	1513
I-type	Low Ni	24	IS 24 -MAG__6_1	304628	3489	109
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__41_1	304452	786	627
I-type	Low Ni	12	IS 12 -MAG 1__36_1	310819	2860	0
I-type	Low Ni	12	IS 12 -MAG 1__41_1	303105	2134	505
I-type	Low Ni	9	IS 9-MAG__2_1	321539	0	10430
I-type	Low Ni	19	IS 19 M__5_1	304980	7650	415
I-type	Low Ni	15	IS 15 C_1	303211	0	0

I-type	Low Ni	22	IS 22 MAG-2__4_1	311591	25815	833
I-type	Low Ni	9	IS 9-MAG__13_1	322906	1351	8302
I-type	Low Ni	11	IS 11-MAG__6_1	307017	22785	1929
I-type	Low Ni	12	IS 12 -MAG 1__11_2	304108	0	386
I-type	Low Ni	17	17NMAG_10	303211	0	359
I-type	Low Ni	11	IS 11-MAG__16_1	306152	4015	1130
I-type	Low Ni	13	IS 13-MAG B__1_1	318856	9258	1787
I-type	Low Ni	9	IS 9-MAG__12_1	305239	11302	259
I-type	Low Ni	19	IS 19 M__20_1	307664	14294	1205
I-type	Low Ni	14	IS 14M Unsifted__26_1	304683	1675	141
I-type	Low Ni	8	IS 8-MAG-2__15_2	312871	11036	2424
I-type	Low Ni	12	IS 12 -MAG 1__22_1	304282	943	0
I-type	Low Ni	12	IS 12 -MAG 1__23_1	304282	943	0
I-type	Low Ni	17	IS 17-MAG__2_1	305218	13267	1465
I-type	Low Ni	24	IS 24 -MAG__3_1	303392	1129	0
I-type	Low Ni	16	IS 16 MAG_2__16_1	301584	62995	0
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__15_1	307017	3900	1068
I-type	Low Ni	8	IS 8-MAG-2__32_2	299878	6331	241
I-type	Low Ni	8	IS 8-MAG-2__24_1	321588	5077	1095
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__58_1	304529	0	898
I-type	Low Ni	20	IS 20 MAG__24_2	307420	4875	614
I-type	Low Ni	8	IS 8-MAG-2__10_2	306838	16907	1744
I-type	Low Ni	20	IS 20 MAG__23_1	303529	578	0
I-type	Low Ni	8	IS 8-MAG-2__13_2	304694	6798	271
I-type	Low Ni	19	IS 19__2_1	306343	2948	1441
I-type	Low Ni	9	IS 9-MAG__8_1	309129	8807	520
I-type	Low Ni	17	IS 17-MAG__13_1	319034	28867	4116
I-type	Low Ni	19	IS 19 M__9_1	302780	4699	62
I-type	Low Ni	7	IS 7 MAG__1_1	309721	14287	799
I-type	Low Ni	20	IS 20 MAG__17_2	302407	0	327
I-type	Low Ni	16	IS 16 MAG_2__6_2	302723	0	0
I-type	Low Ni	12	IS 12-MAG-2__22_1	302679	311	0
I-type	Low Ni	19	19NMAG_66	303100	1518	628
I-type	Low Ni	8	IS 8-MAG-2__12_2	303698	545	398
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__24_1	305375	0	0
I-type	Low Ni	20	IS 20 MAG__26_2	307435	2168	2340
I-type	Low Ni	19	IS 19 M__26_1	303824	547	0
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__21_1	314461	4386	3424
I-type	Low Ni	16	IS 16 MAG_2__19_1	302695	26223	0
I-type	Low Ni	11	IS 11-MAG__3_1	302476	5357	0
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__45_1	304274	3406	292
I-type	Low Ni	8	IS 8-MAG-2__16_1	303874	0	0
I-type	Low Ni	12	IS 12 -MAG 1__30_1	304146	484	551
I-type	Low Ni	13	IS 13-M-1__20_1	309800	2260	709
I-type	Low Ni	12	IS 12 -MAG 1__27_1	315155	6798	9696
I-type	Low Ni	21	IS 21__2_1	305416	532	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__25_1	321377	0	6183

I-type	Low Ni	12	IS 12-MAG-2__21_2	303142	4282	0
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__63_1	309629	0	849
I-type	Low Ni	22	IS 22-MAG -1_right__3_1	302585	0	0
I-type	Low Ni	12	IS 12-MAG-2__17_2	304442	3901	502
I-type	Low Ni	13	IS 13-M-1__17_1	307994	1208	2065
I-type	Low Ni	8	IS 8-MAG-2__27_1	301094	0	635
I-type	Low Ni	22	IS 22-MAG-3__7_1	305239	143	3790
I-type	Low Ni	11	IS 11-MAG__9_1	303248	3760	0
I-type	Low Ni	14	IS 14M Unsifted__9_1	306804	1849	894
I-type	Low Ni	12	IS 12-MAG-2__14_2	308090	9670	757
I-type	Low Ni	14	IS 14M Unsifted__19_1	303889	3449	117
I-type	Low Ni	12	IS 12 -MAG 1__21_2	323199	0	0
I-type	Low Ni	8	IS 8-MAG-2__1_1	302148	2528	0
I-type	Low Ni	4	IS 4-MAG-A__6_2	299707	4128	0
I-type	Low Ni	19	19MAGx_13	301177	0	0
I-type	Low Ni	19	19NMAG_62	301190	0	0
I-type	Low Ni	8	IS 8-MAG-2__26_1	302948	2292	0
I-type	Low Ni	15	IS 15 B_1	302555	0	0
I-type	Low Ni	19	19NMAG_64	300566	2187	454
I-type	Low Ni	4	IS 4-MAG-A__2_2	301636	4612	276
I-type	Low Ni	16	IS 16 MAG_2__5_2	301914	0	0
I-type	Low Ni	19	19NMAG_59	300764	0	59
I-type	Low Ni	19	19MAGx_17	301508	0	0
I-type	Low Ni	19	19NMAG_72	291486	0	0
I-type	Low Ni	19	19MAGx_16	299706	0	0
I-type	Low Ni	8	IS 8-MAG-2__4_2	303195	1078	26
I-type	Low Ni	19	19NMAG_63	301288	558	181
I-type	Low Ni	24	IS 24 -MAG__2_1	302160	8166	0
I-type	Low Ni	12	IS 12 -MAG 1__26_1	301734	0	0
I-type	Low Ni	19	19MAGN_61	301094	0	0
I-type	Low Ni	19	19NMAG_56	301816	0	232
I-type	Low Ni	17	IS 17-MAG__16_1	299372	1412	566
I-type	Low Ni	12	IS 12 -MAG 1__29_1	302833	0	0
I-type	Low Ni	21	IS 21__3_1	302564	0	351
I-type	Low Ni	19	IS 19 M__22_1	303896	7406	0
I-type	Low Ni	19	19MAG_4 part2	301434	0	859
I-type	Low Ni	17	17NMAG_13	302636	0	277
I-type	Low Ni	12	IS 12-MAG-2__15_1	303075	3030	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__45_1	304219	0	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__58_1	303600	0	0
I-type	Low Ni	13	IS 13-M-1__4_1	302291	28908	22
I-type	Low Ni	19	19MAG_2	302995	0	1010
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__41_1	302236	11252	2706
I-type	Low Ni	19	19NMAG_55	300711	0	784
I-type	Low Ni	17	17NMAG_33	303534	0	0
I-type	Low Ni	12	IS 12 S3_1	308739	109	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__50_1	306017	0	0

I-type	Low Ni	13	IS 13M UNSIFTED 1 s__56_1	304250	0	0
I-type	Low Ni	17	17MAG_22	302679	0	148
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__50_1	306219	2032	593
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__19_1	304622	13	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__19_1	305194	0	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__15_1	304430	8872	0
I-type	Low Ni	19	19NMAG_1	302643	0	1387
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__63_1	307685	3607	0
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__17_1	302515	6464	0
I-type	Low Ni	19	19NMAG_45	304570	1049	1697
I-type	Low Ni	13	IS 13M UNSIFTED 60 s__17_1	302010	7842	389
I-type	Low Ni	17	17NMAG_1	304293	831	1338
I-type	Low Ni	15	IS 15 M 2__6_2	303118	45031	0
I-type	Low Ni	19	IS 19 M__25_1	302451	57807	422
I-type	Low Ni	17	17MAG_31	304884	0	808
I-type	Low Ni	17	17MAG_21	307583	0	554
I-type	Low Ni	24	IS 24 -MAG__13_1	307228	664	1791
I-type	Low Ni	17	17MAG_17	314489	332	3192
I-type	Low Ni	17	17MAG_7	314790	431	3272
I-type	Low Ni	9	IS 9-MAG__14_1	306549	6991	915
I-type	Low Ni	15	IS 15 M 2__7_1	307075	18239	721
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__54_1	306314	3743	289
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__61_1	307945	0	0
I-type	Low Ni	17	17MAG_13	306549	842	1335
I-type	Low Ni	17	17MAG_3	306743	1051	1360
I-type	Low Ni	19	19NMAG_60	306014	7436	1930
I-type	Low Ni	17	17NMAG_16	306426	1572	2501
I-type	Low Ni	8	IS 8-MAG-2__7_1	313229	12079	3769
I-type	Low Ni	17	17MAG_4	307099	6483	2845
I-type	Low Ni	17	17MAG_14	307282	6613	2814
I-type	Low Ni	19	19MAG_1	308981	383	2004
I-type	Low Ni	17	17NMAG_3	309058	0	740
I-type	Low Ni	8	IS 8-MAG-2__23_1	314708	4997	1727
I-type	Low Ni	15	IS 15 MAG-1__8_2	305036	23560	799
I-type	Low Ni	19	19MAG_4 part1	308221	0	1102
I-type	Low Ni	17	17NMAG_33 TI	313464	0	0
I-type	Low Ni	16	IS 16 MAG_2__15_1	320752	8234	0
I-type	Low Ni	17	17MAG_45	319207	1275	3687
I-type	Low Ni	14	IS 14M Unsifted__37_1	307241	15108	516
I-type	Low Ni	4	IS 4-MAG-A__14_2	306685	37256	1523
I-type	Low Ni	17	17NMAG_14	315012	0	6850
I-type	Low Ni	13	IS 13M UNSIFTED 1 s__21_1	311553	762	1386
I-type	Low Ni	17	IS 17-MAG__5_2	302101	41437	10841
I-type	Low Ni	8	IS 8-MAG-2__5_1	318119	4581	1624
S-type		13	IS 13M UNSIFTED 60 s__39_1	424509	6207	300999
S-type		13	IS 13M UNSIFTED 1 s__39_1	426929	0	297282
S-type		19	IS 19 M__4_1	423368	7216	284811

S-type	13	IS 13-M-1__15_1	423460	11655	207412
S-type	14	IS 14M Unsifted__16_1	429669	1838	189171
S-type	13	IS 13 NM-1_1	409190	19531	238724
S-type	24	IS 24 -MAG__10_1	414197	5606	231467
S-type	12	IS 12 -MAG 1__4_2	426257	1196	174517
S-type	12	IS 12 -MAG 1__17_1	416818	1655	208220
S-type	13	IS 13M UNSIFTED 1 s__18_1	419059	0	210453
S-type	14	IS 14M Unsifted__17_1	420644	8407	172138
S-type	17	17MAG_25	430676	1001	123697
S-type	13	IS 13M UNSIFTED 1 s__22_1	413807	0	209210
S-type	13	IS 13M UNSIFTED 60 s__18_1	415868	5982	195664
S-type	13	IS 13M UNSIFTED 60 s__22_1	412143	3971	206779
S-type	22	IS 22-MAG -1__6_1	410058	6150	212022
S-type	13	IS 13M UNSIFTED 60 s__60_1	414126	30071	155147
S-type	12	IS 12-MAG-3__1_1	405851	5743	219176
S-type	17	IS 17-MAG__6_1	409178	1878	200909
S-type	19	19NMAG_6	429362	0	119911
S-type	13	IS 13M UNSIFTED 1 s__60_1	415101	18949	153922
S-type	19	19MAGx_9	403826	3268	215969
S-type	13	IS 13M UNSIFTED 1 s__30_1	411375	3511	158050
S-type	13	IS 13M UNSIFTED 60 s__30_1	411003	4453	155165
S-type	13	IS 13-M-1__10_1	409782	8888	162241
S-type	11	IS 11-MAG__2_1	400244	8028	213773
S-type	4	IS 4-MAG-A__13_2	404890	2454	198300
S-type	13	IS 13M UNSIFTED 1 s__2_1	404979	3124	165528
S-type	15	IS 15 MAG-1__9_1	403869	2570	188538
S-type	20	IS 20 MAG__16_1	401715	14025	176990
S-type	19	IS 19 M__2_1	405519	1672	184568
S-type	13	IS 13M UNSIFTED 1 s__33_1	397646	0	221201
S-type	17	17NMAG_7	404935	576	179978
S-type	13	IS 13M UNSIFTED 60 s__33_1	396531	4933	215750
S-type	22	IS 22-MAG -1_right__1_1	402847	1438	170453
S-type	17	17MAG_12	403453	732	173962
S-type	19	IS 19 M__6_1	397711	17093	180988
S-type	12	IS 12 -MAG 1__8_1	396431	2567	196626
S-type	20	IS 20 MAG__14_1	403281	4473	163112
S-type	12	IS 12 -MAG 1__28_1	391487	3917	198416
S-type	24	IS 24 -MAG__8_1	399426	9243	182589
S-type	11	IS 11-MAG-3__3_1	400735	13804	144242
S-type	11	IS 11-MAG__8_1	395308	837	188612
S-type	17	17MAG_2	402816	807	171423
S-type	8	IS 8-MAG-2__2_2	398296	11618	164366
S-type	13	IS 13M UNSIFTED 60 s__64_1	395157	2719	168690
S-type	13	IS 13M UNSIFTED 1 s__64_1	395504	0	171417
S-type	19	IS 19 M__7_1	409863	21459	116216
S-type	22	IS 22-MAG -1__15_1	410799	4097	130001
S-type	12	IS 12-MAG-2__12_2	398339	27823	147629

S-type	13	IS 13M UNSIFTED 60 s__20_1	403221	7352	147499
S-type	13	IS 13M UNSIFTED 1 s__20_1	402884	7077	148604
S-type	13	IS 13-M-1__7_1	393705	5416	194176
S-type	13	IS 13-M-1__13_1	392377	23543	176165
S-type	17	17MAG_1	387780	1159	200283
S-type	17	17MAG_11	387676	787	200340
S-type	16	IS 16 M 1__1_1	397982	23047	147977
S-type	19	19MAGx_10	390505	524	189744
S-type	17	IS 17-MAG__10_1	401322	5121	135113
S-type	19	19NMAG_68 NI	379679	0	162195
S-type	13	IS 13-M-1__2_1	392930	11271	160558
S-type	16	IS 16 MAG_2__11_1	392585	10715	152189
S-type	22	IS 22-MAG -1__4_2	384194	3258	186056
S-type	13	IS 13M UNSIFTED 60 s__10_1	410549	2802	117792
S-type	12	IS 12 -MAG 1__43_1	386998	2951	178097
S-type	12	IS 12 -MAG 1__13_1	391942	1973	181966
S-type	16	IS 16 MAG_2__1_1	390265	6364	186370
S-type	17	IS 17-MAG__14_1	393716	5683	172129
S-type	19	19NMAG_71	390771	0	172736
S-type	18	IS 18-MAG__2_1	387182	709	180884
S-type	13	IS 13M UNSIFTED 60 s__55_1	389798	5660	164231
S-type	14	IS 14M Unsifted__23_1	393024	2170	156739
S-type	19	19MAGx_5	393166	857	164627
S-type	13	IS 13M UNSIFTED 60 s__43_1	402816	12228	127159
S-type	17	17MAG_39	393964	758	157239
S-type	13	IS 13M UNSIFTED 1 s__65_1	393835	0	168839
S-type	12	IS 12 -MAG 1__2_1	390055	3486	185381
S-type	19	19MAGx_18	392870	0	157287
S-type	13	IS 13M UNSIFTED 60 s__34_1	386499	1873	174983
S-type	13	IS 13M UNSIFTED 60 s__65_1	394167	1888	162827
S-type	12	IS 12 -MAG 1__12_2	390458	397	165424
S-type	12	IS 12-MAG-2__1_1	391764	34887	135656
S-type	13	IS 13M UNSIFTED 1 s__34_1	386189	0	174984
S-type	4	IS 4-MAG-A__10_2	387464	18196	153478
S-type	13	IS 13M UNSIFTED 60 s__36_1	386733	4422	170722
S-type	13	IS 13M UNSIFTED 60 s__8_1	396434	129	139751
S-type	13	IS 13M UNSIFTED 1 s__43_1	401711	1298	131719
S-type	13	IS 13M UNSIFTED 60 s__2_1	393091	5716	149331
S-type	17	17NMAG_11 NI	373299	0	164624
S-type	13	IS 13M UNSIFTED 60 s__62_1	392909	8642	154631
S-type	20	IS 20 MAG__18_1	390754	3698	165615
S-type	4	IS 4-MAG-A__1_1	391862	3167	155502
S-type	13	IS 13M UNSIFTED 1 s__55_1	387716	248	156768
S-type	13	IS 13M UNSIFTED 1 s__10_1	408362	3768	101879
S-type	13	IS 13M UNSIFTED 1 s__36_1	386336	9433	162852
S-type	14	IS 14M Unsifted__18_1	378760	15774	176632
S-type	13	IS 13M UNSIFTED 1 s__62_1	394483	0	152458

S-type	13	IS 13M UNSIFTED 60 s__59_1	382106	2619	171259
S-type	22	IS 22-MAG -1__5_1	387769	2100	162890
S-type	13	IS 13M UNSIFTED 1 s__59_1	382423	152	172103
S-type	20	IS 20 MAG__10_2	385922	5603	175092
S-type	13	IS 13M UNSIFTED 1 s__8_1	394525	0	134670
S-type	17	17MAG_5	384722	4435	142980
S-type	17	17NMAG_9	387752	590	159571
S-type	12	IS 12 -MAG 1__16_2	392582	1569	141980
S-type	13	IS 13M UNSIFTED 1 s__57_1	387559	0	161666
S-type	24	IS 24_1	382774	832	166896
S-type	12	IS 12 -MAG 1__18_2	383750	1747	165346
S-type	16	IS 16 MAG_2__18_1	382297	35250	154487
S-type	8	IS 8-MAG-2__8_1	383587	444	163208
S-type	18	IS 18-MAG__4_1	378493	61459	144134
S-type	19	IS 19 M__21_1	380702	8096	165854
S-type	17	17NMAG_23	385013	854	161215
S-type	19	19NMAG_48	382990	1928	173897
S-type	8	IS 8-MAG-2__18_2	380087	20667	159243
S-type	13	IS 13M UNSIFTED 60 s__29_1	387367	0	148170
S-type	22	IS 22-MAG -1__22_2	377266	40443	151334
S-type	13	IS 13-M-1__6_1	378577	6524	171987
S-type	13	IS 13M UNSIFTED 60 s__57_1	384275	8322	154590
S-type	17	17MAG_15	383527	4391	137149
S-type	20	IS 20 MAG__11_1	382461	2744	170687
S-type	11	IS 11-MAG-3__2_2	387805	6333	153330
S-type	17	17NMAG_26	379501	1417	157207
S-type	19	19MAGx_1	380843	2513	163734
S-type	4	IS 4-MAG-A__12_1	381232	3839	157362
S-type	17	17MAG_10	385272	629	161503
S-type	13	IS 13M UNSIFTED 1 s__29_1	386587	0	141851
S-type	16	IS 16 MAG_2__3_1	381647	6672	153113
S-type	17	IS 17-MAG__1_1	387221	3741	148440
S-type	19	19MAGx_2	386114	0	142145
S-type	11	IS 11-MAG__7_1	379822	21810	160793
S-type	17	17MAG_20	384258	551	160477
S-type	13	IS 13M UNSIFTED 60 s__49_1	377028	4001	159079
S-type	24	IS 24 -MAG__9_1	379892	9425	164621
S-type	22	IS 22-MAG-3__5_1	377493	13803	158316
S-type	9	IS 9-MAG__9_1	385431	6131	149577
S-type	16	IS 16 M 1__3_1	379262	18258	146619
S-type	14	IS 14M Unsifted__3_1	374306	13713	157095
S-type	14	IS 14M Unsifted__43_1	374425	15367	159829
S-type	13	IS 13M UNSIFTED 1 s__49_1	377159	677	156463
S-type	17	17NMAG_11	375072	203	159579
S-type	14	IS 14M Unsifted__7_1	385942	3437	140193
S-type	15	IS 15 M 2__5_1	375263	34029	154842
S-type	13	IS 13M UNSIFTED 1 s__13_1	382128	0	158458

S-type	13	IS 13M UNSIFTED 60 s__13_1	381418	4902	153600
S-type	13	IS 13M UNSIFTED 60 s__47_1	377747	1790	151576
S-type	19	19NMAG_47	383577	1718	130369
S-type	17	17MAG_19	378881	141	142437
S-type	19	19MAGx_3	378285	0	146778
S-type	14	IS 14M Unsifted__21_1	378985	775	152943
S-type	22	IS 22-MAG-3__1_1	379542	8372	154466
S-type	13	IS 13M UNSIFTED 60 s__66_1	376510	5122	132240
S-type	22	IS 22-MAG -1__10_1	375415	4217	153202
S-type	12	IS 12-MAG-3__9_2	377946	13580	149391
S-type	22	IS 22-MAG -1__9_1	375055	2969	153340
S-type	19	19NMAG_53	376505	261	152475
S-type	13	IS 13M UNSIFTED 60 s__32_1	379317	11345	131291
S-type	13	IS 13M UNSIFTED 1 s__47_1	376041	0	150669
S-type	4	IS 4-MAG-A__16_1	378887	23471	127728
S-type	12	IS 12-MAG-2__19_2	377050	21534	134525
S-type	17	17NMAG_21	371953	145	148820
S-type	21	IS 21__1_1	373203	1147	150366
S-type	17	17MAG_18	371257	356	149113
S-type	17	17MAG_6 Ni	377094	0	145479
S-type	17	17MAG_9	377741	393	136845
S-type	9	IS 9-MAG__11_1	369549	22070	132957
S-type	13	IS 13M UNSIFTED 60 s__48_1	375086	2004	147777
S-type	24	IS 24 -MAG__12_1	374476	3459	140133
S-type	12	IS 12-MAG-3__2_2	371375	34685	117133
S-type	13	IS 13M UNSIFTED 1 s__48_1	375017	0	148109
S-type	13	IS 13M UNSIFTED 1 s__32_1	379746	1035	130212
S-type	17	17MAG_16	375990	519	142592
S-type	12	IS 12 -MAG 1__35_1	373818	7830	146280
S-type	9	IS 9-MAG__7_1	379971	8114	132614
S-type	17	17MAG_9 Ni	374902	0	138677
S-type	13	IS 13-M-1__3_1	378001	3156	151612
S-type	17	17MAG_6	375637	70	141349
S-type	17	17MAG_43	375030	2547	146452
S-type	4	IS 4-MAG-A__7_2	371096	5212	145881
S-type	14	IS 14M Unsifted__10_1	375986	0	128374
S-type	12	IS 12 -MAG 1__42_1	371517	2773	141660
S-type	13	IS 13-M-1__14_2	377496	23144	122658
S-type	12	IS 12-MAG-2__9_2	368173	20237	141666
S-type	13	IS 13M UNSIFTED 60 s__51_1	373830	8156	136496
S-type	20	IS 20 MAG__1_1	372851	5716	144065
S-type	17	17MAG_8	368718	617	146044
S-type	8	IS 8-MAG-2__30_1	371528	9795	137218
S-type	17	17MAG_19 Ni	373679	0	135439
S-type	19	19MAG_6	378722	0	131988
S-type	13	IS 13M UNSIFTED 1 s__66_1	374396	0	121926
S-type	12	IS 12 -MAG 1__31_1	379710	4441	125749

S-type	12	IS 12-MAG-2__5_2	376986	29016	111321
S-type	11	IS 11-MAG__1_1	373217	11194	131566
S-type	20	IS 20 MAG__4_1	369318	1780	144201
S-type	15	IS 15 MAG-1__2_1	370937	20285	123835
S-type	10	IS 10 -MAG__2_1	372166	39812	102606
S-type	9	IS 9-MAG__5_1	374237	9655	141036
S-type	11	IS 11-MAG__14_1	371673	4709	134151
S-type	8	IS 8-MAG-2__3_2	359089	107679	118941
S-type	8	IS 8-MAG-2__19_2	365953	22196	127730
S-type	12	IS 12 -MAG 1__38_1	366392	7667	128870
S-type	19	19MAGx_14	368669	1372	141865
S-type	17	17MAG_5 Fe	383950	2433	90895
S-type	12	IS 12 -MAG 1__40_1	370727	1653	142155
S-type	14	IS 14M Unsifted__13_1	367134	8334	146020
S-type	19	19NMAG_74	365689	0	128389
S-type	19	19MAGx_6	372389	813	119533
S-type	17	17NMAG_20	370520	286	138816
S-type	17	17MAG_40	403234	3473	54414
S-type	13	IS 13M UNSIFTED 60 s__16_1	381318	3144	104806
S-type	19	19NMAG_49	370319	0	128549
S-type	17	17MAG_26	371036	1648	131322
S-type	12	IS 12-MAG-3__11_1	356946	134119	98393
S-type	13	IS 13M UNSIFTED 1 s__51_1	371212	0	128067
S-type	17	17NMAG_25	367762	882	138793
S-type	15	IS 15 M 2__2_1	365517	27091	123832
S-type	13	IS 13M UNSIFTED 60 s__53_1	371443	9469	114131
S-type	13	IS 13M UNSIFTED 1 s__44_1	368900	0	126508
S-type	13	IS 13M UNSIFTED 1 s__23_1	371047	0	131794
S-type	4	IS 4-MAG-A__9_1	372479	7930	125426
S-type	13	IS 13-M-1__18_1	365438	35118	128515
S-type	13	IS 13M UNSIFTED 60 s__44_1	366644	3114	128126
S-type	12	IS 12-MAG-3__4_1	365468	11868	128423
S-type	19	19NMAG_54	363051	0	132898
S-type	13	IS 13M UNSIFTED 1 s__16_1	378318	0	103359
S-type	13	IS 13M UNSIFTED 60 s__53_1 1	370494	7575	111679
S-type	13	IS 13M UNSIFTED 60 s__53_1 B	370494	7575	111679
S-type	17	17NMAG_19	367483	0	119257
S-type	4	IS 4-MAG-A__11_2	367087	4979	129976
S-type	12	IS 12-MAG-3__7_2	370150	13965	115803
S-type	17	17MAG_5 Cr	350790	0	116990
S-type	4	IS 4-MAG-A__5_2	364901	6516	127519
S-type	13	IS 13M UNSIFTED 60 s__23_1	369828	1289	124191
S-type	15	IS 15 M 2__3_2	377178	0	104067
S-type	4	IS 4-MAG-A__8_2	363858	19925	122336
S-type	15	IS 15 MAG-1__12_2	365983	33703	99723
S-type	19	19NMAG_68	367180	0	104984
S-type	19	IS 19 M__1_1	360764	18503	118894

S-type	19	19NMAG_47-Ni	347055	0	82940
S-type	14	IS 14M Unsifted__45_1	365768	8030	126943
S-type	12	IS 12-MAG-2__7_2	362026	12355	127929
S-type	14	IS 14M Unsifted__34_1	361564	18690	124140
S-type	13	IS 13M UNSIFTED 1 s__53_1	372980	2239	96437
S-type	17	IS 17-MAG__7_2	363980	8002	121013
S-type	12	IS 12 -MAG 1__19_1	365921	0	121278
S-type	17	17NMAG_22	360451	0	116136
S-type	14	IS 14M Unsifted__8_1	366701	812	116550
S-type	4	IS 4-MAG-A__3_2	365103	6435	107461
S-type	17	17NMAG_12	361089	0	117706
S-type	19	IS 19 M 24 Large	364540	49495	99600
S-type	11	IS 11-MAG__12_1	368234	8172	118065
S-type	14	IS 14M Unsifted__5_1	366293	2782	115064
S-type	19	19MAGx_15	370226	520	102140
S-type	14	IS 14M Unsifted__44_1	361730	6849	117624
S-type	17	17NMAG_18 NI	357564	0	113415
S-type	19	19MAGx_12	361581	0	120731
S-type	14	IS 14M Unsifted__39_1	367746	1259	91307
S-type	13	IS 13-M-1__12_1	366626	3604	114981
S-type	17	17NMAG_26 NI	286392	0	89431
S-type	13	IS 13M UNSIFTED 1 s__5_1	360880	11082	112295
S-type	13	IS 13M UNSIFTED 60 s__5_1	362338	1554	113247
S-type	9	IS 9-MAG__10_1	370136	17673	58171
S-type	19	IS 19 M__17_1	360662	28557	90148
S-type	11	IS 11-MAG__5_1	366136	7504	103076
S-type	15	IS 15 MAG-1__6_1	363257	26980	91305
S-type	13	IS 13M UNSIFTED 1 s__31_1	357287	0	100716
S-type	12	IS 12 -MAG 1__6_2	363394	0	97661
S-type	17	IS 17-MAG__11_1	355538	38881	102306
S-type	12	IS 12 -MAG 1__37_1	363015	10870	99312
S-type	14	IS 14M Unsifted__12_1	356571	3799	113531
S-type	14	IS 14M Unsifted__41_1	357084	14762	113940
S-type	12	IS 12 -MAG 1__24_1	378208	3	57538
S-type	9	IS 9B S15	354458	805	113525
S-type	12	IS 12-MAG-2__13_1	349418	44344	102832
S-type	22	IS 22 MAG-2__1_1	357572	21747	92802
G-type	13	IS 13M UNSIFTED 60 s__31_1	355351	554	90465
G-type	8	IS 8-MAG-2__11_1	357208	8474	93016
G-type	19	IS 19 M__3_1	360493	30864	77353
G-type	19	19MAGx_22	344059	0	109472
G-type	15	IS 15 M 2__4_1	353694	55476	78226
G-type	12	IS 12 B_1	368411	823	63222
G-type	16	IS 16 MAG_2__12_1	357962	28965	81501
G-type	11	IS 11-MAG__11_1	364730	8567	73700
G-type	12	IS 12-MAG-2__20_1	350310	21238	99386
G-type	17	17NMAG_22 NI	336269	0	89668

G-type	23	IS 23 M__1_1	352447	9314	74961
G-type	17	17NMAG_18	354958	0	96309
G-type	14	IS 14M Unsifted__31_1	357474	11227	87198
G-type	16	IS 16 M 1__4_2	353641	4336	95788
G-type	13	IS 13M UNSIFTED 1 s__52_1	343901	0	96053
G-type	12	IS 12-MAG-2__16_1	348245	8453	92979
G-type	22	IS 22-MAG -1__21_2	358786	5275	72340
G-type	15	IS 15 MAG-1__1_1	355498	20718	89431
G-type	13	IS 13M UNSIFTED 60 s__11_1	350645	558	79357
G-type	17	IS 17-MAG__8_2	361608	47022	42746
G-type	13	IS 13M UNSIFTED 1 s__11_1	353094	0	72131
G-type	22	IS 22-MAG -1_right__2_1	353619	0	82844
G-type	20	IS 20 MAG__20_1	350246	5790	85655
G-type	13	IS 13M UNSIFTED 60 s__52_1	341358	0	88820
G-type	12	IS 12-MAG-2__11_1	346400	19129	86667
G-type	17	17NMAG_27	348925	0	76819
G-type	19	IS 19 M__19_1	362724	4658	47359
G-type	14	IS 14M Unsifted__28_1	349334	3676	78842
G-type	12	IS 12 -MAG 1__34_1	352140	15203	71386
G-type	20	IS 20 MAG__21_1	354097	967	67521
G-type	14	IS 14M Unsifted__20_1	358630	0	63086
G-type	19	19MAGx_11	359099	0	49463
G-type	9	IS 9-MAG__4_1	339786	4022	78928
G-type	16	IS 16 M 1__2_1	354153	21570	59244
G-type	14	IS 14M Unsifted__14_1	352159	88	83659
G-type	13	IS 13-MAG B__4_1	345516	18886	63884
G-type	11	IS 11-MAG__15_1	336416	6496	95302
G-type	12	IS 12-MAG-2__18_2	342171	24157	88719
G-type	19	19MAGx_8	345873	0	58888
G-type	14	IS 14M UNSIFTED B 60 s_8_1	336664	28680	74511
G-type	17	IS 17-MAG__3_1	351304	0	52913
G-type	12	IS 12-MAG-3__10_1	343264	15493	54634
G-type	19	19NMAG_51	345182	0	49683
G-type	24	IS 24_2	351608	762	56158
G-type	12	IS 12-MAG-2__4_2	355387	23962	38717
G-type	19	19MAGx_19	344882	0	46529
G-type	17	17MAG_40 Cu	357525	0	31244
G-type	13	IS 13M UNSIFTED 1 s__42_1	327446	0	67877
G-type	13	IS 13M UNSIFTED 60 s__42_1	326083	145	68734
G-type	14	IS 14M Unsifted__35_1	335228	1216	51206
G-type	17	17MAG_29	342864	249	48349
G-type	24	IS 24 -MAG__5_1	340260	10581	35862
G-type	13	IS 13-M-1__5_1	337168	4734	57868
G-type	14	IS 14M Unsifted__46_1	338658	0	41353
G-type	17	17MAG_38 Ni 2	323440	0	42638
G-type	17	17MAG_38 Ni 1	323342	0	36850
G-type	19	19MAGx_23	322806	0	43618

G-type		12	IS 12-MAG-2__3_1	328044	28748	34990
G-type		8	IS 8-MAG-2__17_2	335003	6231	27975
G-type		9	IS 9-MAG__6_1	322387	12044	25626
G-type		8	IS 8-MAG-2__31_2	338208	11794	14422
G-type		13	IS 13-MAG B__6_2	320600	9724	31562
G-type		17	IS 17-MAG__12_2	326308	4669	23151
G-type		14	IS 14M Unsifted__40_1	316210	0	29934
G-type		17	17MAG_38	318704	0	24428
G-type		23	IS 23 M__2_1	315572	0	27717
G-type		19	19MAGx_26	313495	0	25064
G-type		8	IS 8-MAG-2__25_1	314169	2042	26197
G-type		12	IS 12 -MAG 1__25_2	316372	0	18807
G-type		17	17MAG_30	332538	2223	15316
G-type		8	IS 8-MAG-2__28_1	313121	4155	19557
G-type		15	IS 15 MAG-1__11_1	326571	12153	12512
G-type		13	IS 13-MAG B__7_1	313702	1119	22777
G-type		22	IS 22-MAG -1__17_1	321809	3568	14547
G-type		19	IS 19__1_1	316606	6058	9016
G-type		20	IS 20 MAG__3_1	302106	0	19229
G-type		12	IS 12 -MAG 1__20_1	310861	1587	16986
G-type		13	IS 13M UNSIFTED 1 s__9_1	320215	0	7847
D-type	High Sr, high Si	17	IS 17-MAG__17_2	433886	9357	8953
D-type	High Sr, high Si	13	IS 13M UNSIFTED 1 s__3_1	465744	14498	5220
D-type	High Sr, high Si	15	IS 15 MAG-1__10_1	438607	33752	3585
D-type	High Sr, high Si	22	IS 22-MAG -1__3_2	452383	19390	4328
D-type	High Sr, high Si	22	IS 22 MAG-2__8_1	426804	25964	6328
D-type	High Sr, high Si	13	IS 13M UNSIFTED 60 s__3_1	463100	18292	10580
D-type	High Sr, high Si	13	IS 13M UNSIFTED 1 s__1_1	433442	12348	7057
D-type	High Sr, high Si	13	IS 13M UNSIFTED 60 s__1_1	432651	22002	12455
D-type	High Sr, high Si	11	IS 11-MAG 2__1_1	437156	24059	10306
D-type	High Sr, high Si	17	17MAG_33 Ni	457285	14467	8073
D-type	High Sr, high Si	19	IS 19 M__14_1	451431	35573	3000
D-type	High Sr, high Si	17	17MAG_36	424408	22240	12236
D-type	High Sr, high Si	20	IS 20 MAG__7_2	412896	33192	10642
D-type	High Sr, high Si	22	IS 22-MAG -1__7_1	419925	27766	11242
D-type	High Sr, high Si	22	IS 22-MAG -1__8_1	419925	27766	11242
D-type	High Sr, high Si	11	IS 11-MAG 2__3_1	443060	21195	10888
D-type	High Sr, high Si	17	17NMAG_29	428520	22656	14645
D-type	High Sr, high Si	17	17NMAG_5	449883	17085	5571
D-type	High Sr, high Si	16	IS 16 MAG_2__4_1	438108	15952	67785
D-type	High Sr, high Si	22	IS 22 MAG-2__6_1	436079	11419	27254
D-type	High Sr, low Si	8	IS 8-MAG-2__9_2	421225	3465	2355
D-type	High Sr, low Si	19	IS 19 M__16_1	408695	1778	8536
D-type	High Sr, low Si	22	IS 22-MAG -1__20_1	397243	7029	9384
D-type	High Sr, low Si	17	17MAG_35 Ni 1	403346	0	714
D-type	High Sr, low Si	17	IS 17-MAG__15_1	391021	39726	5580
D-type	High Sr, low Si	24	IS 24 -MAG__11_1	392147	15882	12288

D-type	High Sr, low Si	17	17MAG_35 Ni 2	391872	0	838
D-type	High Sr, low Si	19	19NMAG_50	378417	2222	12512
D-type	High Sr, low Si	17	17MAG_35	372911	2299	4300
D-type	High Sr, low Si	19	IS 19 M__18_1	369510	2982	8808
D-type	High Sr, low Si	22	IS 22-MAG-3__4_2	362544	97144	14232
D-type	High Sr, low Si	17	17MAG_33	350941	0	9317
D-type	High Sr, low Si	17	17NMAG_6	353270	2353	7556
D-type	High Sr, low Si	9	IS 9-MAG__3_1	345775	21453	16793
D-type	High Sr, low Si	19	19MAGx_4	347850	9926	4935
D-type	High Sr, low Si	14	IS 14M UNSIFTED B 60 s_2_1	333734	60742	5309
D-type	High Sr, low Si	14	IS 14M Unsifted__48_1	344885	7822	11025
D-type	High Sr, low Si	17	17NMAG_2	342753	1226	10054
D-type	High Sr, low Si	15	IS 15 M 2__8_1	328989	73345	2554
D-type	High Sr, low Si	13	IS 13-MAG B__3_2	314504	139827	1078
D-type	High Sr, low Si	17	17NMAG_28	329648	0	628
D-type	High Sr, low Si	22	IS 22-MAG -1__16_1	326467	4293	0
D-type	High Sr, low Si	14	IS 14M Unsifted__47_1	329013	5190	2721
D-type	Low Sr, high Si	13	IS 13 NM-2_1	452594	32227	4936
D-type	Low Sr, high Si	14	IS 14 NM__1_1	470732	32625	1735
D-type	Low Sr, high Si	17	IS 17-MAG__19_1	463670	32183	2658
D-type	Low Sr, high Si	9	IS 9-NM_1	466003	31583	8480
D-type	Low Sr, high Si	16	IS 16 MAG_2__14_1	447115	60422	1522
D-type	Low Sr, high Si	17	17NMAG_4	460671	26930	9084
D-type	Low Sr, high Si	20	IS 20 MAG__8_2	467580	28765	806
D-type	Low Sr, high Si	12	IS 12 M4_1	461701	40861	16899
D-type	Low Sr, high Si	13	IS 13 NM-1_2	447871	50816	9366
D-type	Low Sr, high Si	22	IS 22 MAG-2__7_1	444757	65937	8190
D-type	Low Sr, high Si	14	IS 14M Unsifted__4_1	464441	27797	10371
D-type	Low Sr, high Si	12	IS 12-MAG-3__6_1	446815	59013	5124
D-type	Low Sr, high Si	19	19MAG_10Fe	449248	27457	19308
D-type	Low Sr, high Si	17	17MAG_28 Particle 2	449969	27637	9556
D-type	Low Sr, high Si	17	17MAG_23	453317	29122	1411
D-type	Low Sr, high Si	14	IS 14M UNSIFTED B 60 s_3_1	446133	30597	11779
D-type	Low Sr, high Si	11	IS 11-MAG 2__2_2	445398	30319	12256
D-type	Low Sr, high Si	12	IS 12-MAG-3__12_1	435658	48719	8952
D-type	Low Sr, high Si	19	19NMAG_7	445505	26015	1709
D-type	Low Sr, high Si	19	19NMAG_8	443276	14831	40910
D-type	Low Sr, high Si	22	IS 22-MAG -1__2_2	440110	34934	923
D-type	Low Sr, high Si	17	IS 17-MAG__18_2	447091	7253	2861
D-type	Low Sr, high Si	20	IS 20 MAG__9_1	438556	28914	15167
D-type	Low Sr, high Si	17	17MAG_27	439219	24942	4035
D-type	Low Sr, high Si	19	IS 19__3_1	438905	34504	3011
D-type	Low Sr, high Si	15	IS 15 M_1	435337	20077	39668
D-type	Low Sr, high Si	14	IS 14M UNSIFTED B 60 s_4_1	432868	42545	10262
D-type	Low Sr, high Si	19	19MAG_5	447848	26489	33684
D-type	Low Sr, high Si	14	IS 14 NM__4_1	425715	33699	23071
D-type	Low Sr, high Si	22	IS 22-MAG -1__19_1	427420	25527	1653

D-type	Low Sr, high Si	14	IS 14M UNSIFTED B 60 s_6_1	424739	27994	2604
D-type	Low Sr, high Si	19	19NMAG_5	437566	9007	3059
D-type	Low Sr, high Si	19	19NMAG_8 Ti	430215	4446	0
D-type	Low Sr, high Si	17	17NMAG_32	439629	3206	1869
D-type	Low Sr, high Si	19	19NMAG_57	424905	20028	2145
D-type	Low Sr, high Si	19	19NMAG_4	422780	17768	0
D-type	Low Sr, high Si	19	19NMAG_7 Ti	418183	26951	0
D-type	Low Sr, high Si	22	IS 22-MAG-3__10_1	411095	50410	2493
D-type	Low Sr, high Si	13	IS 13M UNSIFTED 60 s__38_1	414005	21392	1305
D-type	Low Sr, high Si	20	IS 20 MAG__5_1	419252	22982	39
D-type	Low Sr, high Si	18	IS 18-MAG__1_1	402636	69149	485
D-type	Low Sr, high Si	20	IS 20 MAG__28_2	415320	35576	852
D-type	Low Sr, high Si	13	IS 13M UNSIFTED 60 s__28_1	417000	24028	1690
D-type	Low Sr, high Si	16	IS 16 MAG_2__10_2	415191	36571	8497
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 1 s__28_1	416861	14098	0
D-type	Low Sr, low Si	17	17MAG_28 Particle 1	412894	24553	1444
D-type	Low Sr, low Si	17	17NMAG_31	414250	19824	3142
D-type	Low Sr, low Si	19	19MAGx_7	412082	20789	2813
D-type	Low Sr, low Si	22	IS 22-MAG -1__13_1	411523	23792	886
D-type	Low Sr, low Si	17	17NMAG_8	406689	18030	1417
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 1 s__7_1	419069	2084	14009
D-type	Low Sr, low Si	19	19MAGx_21	408206	26557	5491
D-type	Low Sr, low Si	22	IS 22 MAG-2__3_1	399739	57659	3157
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 60 s__7_1	419916	5911	21699
D-type	Low Sr, low Si	14	IS 14M UNSIFTED B 60 s_7_1	403493	41715	1954
D-type	Low Sr, low Si	12	IS 12 -MAG 1__7_1	406772	20366	1614
D-type	Low Sr, low Si	20	IS 20 MAG__25_2	404791	20375	2106
D-type	Low Sr, low Si	11	IS 11-MAG__10_1	406264	25269	2098
D-type	Low Sr, low Si	11	IS 11-MAG__4_1	405621	21909	1313
D-type	Low Sr, low Si	19	IS 19 M__12_1	405510	40967	1214
D-type	Low Sr, low Si	19	19NMAG_58	399114	20565	2051
D-type	Low Sr, low Si	19	19NMAG_73	400576	18561	0
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 1 s__6_1	398404	22678	0
D-type	Low Sr, low Si	4	IS 4-MAG-A__4_2	394778	56956	3457
D-type	Low Sr, low Si	19	19NMAG_70	399980	14959	0
D-type	Low Sr, low Si	17	17MAG_32	397453	23316	3580
D-type	Low Sr, low Si	13	IS 13-M-1__16_2	398270	22080	521
D-type	Low Sr, low Si	22	IS 22-MAG -1__8_2	401027	21054	3138
D-type	Low Sr, low Si	17	17MAG_25 Ti	400172	14789	9836
D-type	Low Sr, low Si	14	IS 14M Unsifted__36_1	400206	22343	2486
D-type	Low Sr, low Si	20	IS 20 MAG__27_1	392043	39113	1648
D-type	Low Sr, low Si	14	IS 14 NM__2_1	395909	8327	12663
D-type	Low Sr, low Si	9	IS 9-MAG__1_1	404750	18563	13320
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 60 s__6_1	396008	22957	3744
D-type	Low Sr, low Si	17	17MAG_24	393962	15312	2811
D-type	Low Sr, low Si	18	IS 18-MAG__3_1	382774	78154	1176
D-type	Low Sr, low Si	8	IS 8-MAG-2__21_2	392537	29683	1718

D-type	Low Sr, low Si	19	19MAGx_20	393339	18436	2908
D-type	Low Sr, low Si	12	IS 12-MAG-2__6_1	383553	61479	1001
D-type	Low Sr, low Si	14	IS 14M Unsifted__24_1	393662	15875	8515
D-type	Low Sr, low Si	19	19MAGx_29	388524	27767	5394
D-type	Low Sr, low Si	19	19MAGx_27	386662	10765	7706
D-type	Low Sr, low Si	17	IS 17-MAG__9_2	386894	4312	338
D-type	Low Sr, low Si	19	19MAGx_24	383297	11477	3843
D-type	Low Sr, low Si	18	IS 18-MAG__6_1	389336	13246	14783
D-type	Low Sr, low Si	21	IS 21__4_1	385227	10297	5105
D-type	Low Sr, low Si	12	IS 12-MAG-3__5_1	388345	14847	21257
D-type	Low Sr, low Si	14	IS 14M UNSIFTED B 60 s_1_1	376127	2570	11873
D-type	Low Sr, low Si	14	IS 14M Unsifted__30_1	380882	7679	4580
D-type	Low Sr, low Si	19	IS 19 M__12_map	372836	27122	355
D-type	Low Sr, low Si	17	17MAG_34	366157	13289	2511
D-type	Low Sr, low Si	19	19MAG_10Ti	366554	7538	15633
D-type	Low Sr, low Si	19	19MAG_67	370964	10404	1620
D-type	Low Sr, low Si	16	IS 16 MAG_2__21_2	360940	6541	121
D-type	Low Sr, low Si	19	19NMAG_67	364132	6792	1173
D-type	Low Sr, low Si	14	IS 14M Unsifted__2_1	349902	38779	9163
D-type	Low Sr, low Si	13	IS 13-M-1__19_1	355786	14855	1615
D-type	Low Sr, low Si	11	IS 11-MAG-3__4_1	348510	860	1573
D-type	Low Sr, low Si	22	IS 22-MAG-3__3_1	344368	59112	19033
D-type	Low Sr, low Si	7	IS 7 MAG__3_1	340175	23496	612
D-type	Low Sr, low Si	19	19MAG_9	339764	8018	7261
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 60 s__46_1	348450	19765	11880
D-type	Low Sr, low Si	19	19NMAG_52	338553	3423	3769
D-type	Low Sr, low Si	19	19MAG_7	336641	4860	1610
D-type	Low Sr, low Si	19	19NMAG_9	338218	5487	4981
D-type	Low Sr, low Si	16	IS 16 MAG_2__20_1	334096	2340	1069
D-type	Low Sr, low Si	7	IS 7 MAG__2_1	333451	7381	1075
D-type	Low Sr, low Si	19	19MAGx_28	328460	2624	1811
D-type	Low Sr, low Si	19	19NMAG_11 FeTi	331773	0	9939
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 60 s__9_1	328940	3775	5641
D-type	Low Sr, low Si	22	IS 22-MAG-3__2_2	331799	16683	4961
D-type	Low Sr, low Si	20	IS 20 MAG__15_2	326975	5818	4681
D-type	Low Sr, low Si	13	IS 13-M-1__8_1	318026	29002	2396
D-type	Low Sr, low Si	8	IS 8-MAG-2__22_1	327176	10512	6195
D-type	Low Sr, low Si	19	19MAG_8	329135	2548	4839
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 60 s__4_1	318945	25125	1849
D-type	Low Sr, low Si	13	IS 13M UNSIFTED 1 s__4_1	319143	14591	545
D-type	Low Sr, low Si	17	17MAG_41	321451	0	4572
D-type	Low Sr, low Si	19	19NMAG_69	310790	5541	4129

tted in Figures 5, 6, 7, 8, 9a, 10a).

Al	Si	K	Ca	Ti	Cr	Mn	Fe	Ni
1792	9531	292	393	2550	99	0	590334	102811
2877	16985	87	999	652	3726	561	601403	51004
1926	6383	0	496	137	36	0	655832	33693
1185	6420	804	374	265	598	0	666758	20804
2637	15437	673	1003	445	718	240	634589	17641
2637	15437	673	1003	445	718	240	634589	17641
624	5442	0	213	0	563	0	677360	16948
1847	10361	186	856	1666	759	0	666997	15516
3719	11186	221	1102	1670	837	0	658125	14888
410	5186	0	177	68	273	0	679923	13882
951	7967	160	2503	184	1431	0	665329	11949
1473	9777	0	365	222	231	122	668334	11938
2839	16898	992	3239	623	2567	1494	642196	10364
2155	8349	128	158	616	1133	0	672229	10233
2555	10085	526	620	884	12623	0	646029	6125
508	5420	0	65	92	708	77	681217	6106
1024	6685	27	188	209	1546	5	682446	4887
7120	30379	1805	11687	1838	205	1489	554693	4058
978	7511	0	1732	139	4068	4394	662231	2418
2079	10805	2401	919	142	331	0	653294	1796
1606	13656	246	411	138	2252	0	654595	1687
1706	6549	38	325	226	1638	0	683447	1589
1791	14091	57	2569	192	90	6366	658832	1320
1120	5503	0	328	104	158	0	687880	1316
1843	11967	129	803	173	317	0	673627	1137
1415	6493	138	190	76	233	2426	682891	1081
794	17454	306	545	227	891	1192	659518	915
2498	8245	145	1542	351	1298	0	676592	816
6329	32498	632	2945	975	483	140	627455	816
3262	11958	226	12819	2885	761	5688	643343	732
2341	16223	2	1302	112	558	1122	649351	605
3793	10237	50	1058	220	210	2901	665113	568
847	7377	11	234	4441	112	2292	674951	547
1395	12584	55	428	93	969	1547	667970	540
915	7942	98	60	128	95	4568	659311	537
7707	19996	938	1208	31811	282	2844	600114	534
518	5748	267	347	39	323	0	687970	510
1539	8808	27	335	289	264	593	679787	505
9970	14269	263	504	25087	1372	3230	621845	494
1808	15759	354	672	279	631	66	654249	472
1890	12763	217	1033	629	816	3017	665313	468
1383	12377	72	1087	3733	276	4022	660153	465
9817	24537	517	2608	22454	966	4609	602164	446
3793	7391	93	1481	1191	156	0	678054	427
726	20042	608	5017	2508	47	608	647100	424

18058	11114	0	874	25932	578	802	618061	420
447	6905	0	34	238	1328	4497	678371	410
3489	10338	342	1457	13804	501	5019	652710	376
1073	7715	281	182	41	436	846	676957	350
871	14497	46	3685	238	172	119	666231	323
1604	18664	174	972	2280	21705	20006	624080	307
3382	16071	71	2268	3737	220	2332	658105	305
426	6361	136	63	13	4111	1533	682733	296
1206	9732	240	213	249	85	4593	649902	296
3659	12586	162	1006	1110	543	3848	665451	288
17933	11456	121	1263	25461	673	943	607251	270
1144	11636	90	214	59	395	9468	659813	262
10306	9058	128	564	24292	2376	2137	627596	256
1451	10614	158	503	305	354	2671	673899	253
1156	13236	0	820	676	624	2692	655970	251
939	11547	399	393	455	1298	3994	663502	248
11203	8818	63	473	24846	2133	1966	631000	231
579	8202	0	96	87	393	3800	681061	230
2853	13103	441	992	182	1511	2762	598971	203
17236	13920	440	746	31216	664	1168	559293	201
35563	19244	421	1397	25826	1540	2538	579229	185
14759	20634	521	281	47202	470	2188	506477	183
2082	10715	37	681	486	750	2245	671971	182
4573	14728	242	5982	375	525	1391	660158	182
1547	8584	41	926	6541	287	3759	670251	174
448	6019	625	65	59	292	2588	683452	172
2955	15408	117	1320	216	262	2767	658666	170
2246	13974	665	2511	735	277	3281	647184	169
1633	12669	864	732	579	763	5958	663118	167
2171	10735	0	1611	167	647	1701	664564	166
3744	8017	0	1384	992	169	0	679519	164
3358	11046	160	5980	43397	492	17063	581721	163
389	15658	559	853	1909	233	1225	661739	152
1928	7710	26	695	125	839	2507	678439	151
1061	7989	429	462	666	319	1666	617017	150
1356	23619	2505	541	1526	197	2658	627321	150
893	9185	43	525	239	535	2702	676979	147
5863	18387	274	4105	1473	468	995	648559	146
2836	19999	372	967	21748	107	1124	607403	144
7472	21961	360	1179	134	146	1791	622053	136
1126	8652	105	431	1401	209	1661	677453	123
1864	8732	0	548	99	800	1695	679202	116
1058	29584	98	195	126	3057	10556	639157	110
695	6873	0	248	155	666	1911	682707	110
16760	15850	215	2421	43758	974	3276	583099	102
2012	12355	150	1126	68	461	5715	663986	99
637	9080	39	121	16	345	5841	680038	90

6181	20114	427	4154	2893	322	1618	623870	86
24939	17941	485	581	36155	1794	3198	580670	83
2072	12991	324	773	1502	371	2133	644995	82
576	8148	35	80	172	390	1258	683475	77
1593	8148	0	538	509	398	32	678019	76
2846	10729	67	953	1008	196	2909	668977	75
16337	26660	1760	3601	2757	124	114	616282	74
819	10364	67	353	4502	225	2173	663014	73
2653	12145	120	4332	21034	315	6731	627591	70
994	9718	133	486	1332	225	2781	676215	69
1678	21771	514	4753	10236	159	1775	630604	63
455	6684	0	100	3446	371	1366	679378	62
455	6684	0	100	3446	371	1366	679378	62
3934	13544	67	1685	249	273	8964	649995	61
546	6242	167	102	1199	143	297	685373	60
1292	9195	243	336	908	208	1606	620845	60
2225	13887	0	634	402	175	3346	665241	60
505	6401	35	451	379	1511	8767	673062	58
8355	29941	304	23094	25926	1537	42	581005	49
1172	8915	115	688	395	138	2759	679760	49
5991	14131	745	1302	1213	587	4733	656121	49
1463	14313	7	1211	112	332	2014	653961	49
621	6059	28	132	1466	336	756	686090	46
1312	9126	93	151	4	178	305	676015	42
1018	10910	51	591	994	264	367	674720	41
2549	19337	166	351	0	308	3508	654479	41
30186	23736	1118	7614	2295	221	5716	573767	39
815	7618	35	170	1810	4210	2125	673832	38
4927	17390	597	769	189	292	2638	646248	37
634	5789	0	80	48	371	3279	686243	37
632	6816	42	62	536	471	3465	684160	36
498	5804	85	190	16	1274	1431	686869	35
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6822	7171	111	55	15218	1564	3232	652721	32
1241	6737	223	423	630	327	1288	684020	30
6140	24995	620	1271	273	500	359	641647	27
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1947	17302	70	288	12939	83	5804	647669	22
10101	8014	689	475	47023	513	3908	596155	20
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28534	15298	253	1442	26974	1454	2083	595358	19

803	7693	172	68	563	603	3371	678238	17
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332	5130	0	48	48	332	1032	689548	15
1166	8056	151	539	39	275	1239	678651	12
3827	19150	479	1803	1543	195	11289	649227	11
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3735	10313	66	1876	20534	1512	9055	634789	8
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3046	12585	695	2030	11660	263	2560	646063	5
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587	5881	32	417	58	578	3604	683766	2
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412	5287	0	85	86	232	1655	684679	0
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756	5037	0	31	54	367	7242	677025	0
1778	5288	113	1084	116	117	5321	677961	0
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618	9420	0	537	1786	304	1766	669164	0
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2682	10031	42	1117	179	315	3372	668390	0
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9351	12610	243	1145	637	140	175	659314	0
2794	11602	119	1578	804	246	611	671880	0
29755	9147	0	2042	7344	153	1879	624304	0
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3218	12905	507	3187	940	160	3359	660162	0
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7038	168159	368	10828	819	4406	2311	177471	1406
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65594	238947	18526	50672	4865	19	2004	123347	289
54840	265851	9015	24812	2937	0	8114	157262	0
40233	256087	5403	117922	3433	27	3946	70095	0
49572	232029	7256	9672	73849	44	1599	148589	0
70766	259698	23349	11318	5281	300	4568	164851	70
71780	239061	21097	47554	5357	28	2812	127774	33
65221	243219	5882	43623	5802	149	1792	161597	0
67761	242776	17650	21691	4876	70	1761	165222	32
51538	235751	8148	117011	3096	33	1633	86915	133
55043	232979	10621	35080	10200	115	1544	166764	10
65347	243549	1109	26182	20512	70	2575	128784	0
56357	214333	8097	94183	10579	102	1960	129336	32
49621	226455	8588	34512	15420	104	1969	207834	24

51759	220408	9569	44385	10600	221	2331	203152	144
90562	221495	12276	10493	6850	129	87	205255	0
45431	231396	6173	28215	19667	0	1927	228958	0
153199	181860	814	2960	12069	81	330	194570	61
50165	222103	5596	9264	11544	16	1157	246475	0
45518	217524	5771	12986	19537	22	2072	249747	0
47701	209457	7542	13078	19911	19	2260	249007	0
39528	201559	10003	23752	12558	109	2353	243112	579
41652	187653	6105	25455	61923	80	2097	236798	9
43143	211936	4258	6298	16447	296	2333	272163	0
39014	186418	5526	35294	17818	122	2052	240378	0
43411	203840	5014	8569	18696	144	2022	265653	0
45659	205689	5562	11266	18336	136	1997	266475	0
63526	191512	10264	23683	14750	272	1420	233523	80
44221	204674	5285	11569	19816	136	1857	280377	0
44482	198379	5154	12590	19467	273	2495	272117	0
45383	198062	4459	12611	23348	41	2261	270744	0
45216	195019	4827	20496	20549	0	1855	269794	0
42933	194985	4580	18916	21313	152	2119	278143	0
40429	190852	6902	28424	13603	9	3230	283257	0
88132	183506	4086	27495	2857	29	6733	250995	0
58241	179728	7763	13474	22245	0	2824	270025	0
48754	172164	12715	19128	18787	316	1576	264350	36
90949	181720	4172	27066	2709	108	6607	237687	18
40373	181791	4991	12933	17386	445	2744	289289	84
41967	184723	4195	13932	22610	189	2315	300236	0
41170	182391	5437	24295	20233	129	2087	295874	124
51778	178859	9053	5789	21245	226	1884	296926	0
41484	182386	3657	8944	23705	365	2282	307089	346
101080	153201	6246	19706	3963	56	0	265154	72
41642	164699	4336	40260	22803	0	2553	289160	0
39994	174556	3680	15354	22770	0	2233	316353	0
31743	176482	4050	20305	20085	257	3511	320129	95
42918	165292	4376	16790	13596	211	1343	299194	0
39874	175652	3336	9322	16657	0	1980	332878	0
41330	166996	11078	7313	29484	0	4179	309557	0
34485	174726	5294	9838	19334	161	2876	331387	0
41718	174138	2937	8968	17921	772	1592	325160	0
33408	175998	4834	7601	23796	127	3292	318580	0
39800	174594	3438	6602	13752	882	844	334319	3
38632	162222	10478	9351	24602	167	5812	314350	145
11854	189830	23647	7920	1755	113	1380	345231	35
57906	167028	1457	18939	16389	352	1119	299179	208
31560	169386	3418	18028	22259	265	3022	326049	322
37749	158095	5126	31627	31419	0	2960	312412	0
31274	148816	5320	4396	24456	186	2485	318735	536
37250	157924	2149	5815	25790	95	3033	342598	0

37260	160047	2572	4532	26132	22	2879	346159	0
31965	151623	4688	7803	10945	321	1210	344272	0
75475	143183	20686	4583	12706	374	2439	320136	1665
61521	122283	1920	20564	70626	0	2045	289992	0
43740	144475	1657	11018	17624	247	1782	365691	40
58987	142854	2256	14001	926	40	8202	379708	0
37838	142784	2454	3919	18717	143	1389	384339	632
48686	140699	1004	16485	15055	904	953	357031	557
70453	128276	8398	32305	4575	96	227	352016	0
54387	137222	2712	18796	5083	805	1022	354438	166
21149	140159	18724	3648	1785	175	14	421431	361
86866	104082	502	43067	11828	354	1413	357177	15
75985	101800	4663	16356	3650	58	0	393682	619
29095	105930	1937	7938	31575	0	3522	429155	0
24401	106790	3111	24402	33534	8	4059	405538	0
39558	91492	741	43278	23184	19	2459	390142	0
16694	113122	2873	1703	627	90	499	495276	95
35569	85038	611	30604	23546	16	2590	428432	0
8457	96270	12888	1656	712	152	276	476116	1265
56223	77290	3891	10624	1720	135	0	477030	21
24357	81785	647	20319	2569	128	0	517191	279
16552	74087	616	9119	22114	296	3932	448719	159
11922	75767	3074	8433	2385	63	338	532246	13
30285	63886	1341	4605	38970	414	2640	470627	232
25954	63658	1247	5984	50794	533	1276	467243	2049
25506	64229	5254	4846	3461	164	572	543262	0
18972	64633	507	10262	378	151	0	555198	0
16729	57160	485	1624	34034	97	4152	531319	0
18228	56791	2129	3506	2989	202	263	576705	3
19079	55067	829	16873	4211	212	460	559553	0
7433	51561	1568	3824	13473	1088	3106	570256	271
12304	44545	434	8862	38000	14	3717	542751	0
12791	39672	583	51116	40899	506	3917	510734	0
22306	36520	1777	1242	55677	425	3298	519419	0
12584	34937	477	9260	39561	282	3852	559530	100
8589	33478	1594	41569	11815	750	3514	547862	78
13696	34282	852	9727	36107	702	1428	558088	179
12672	33441	461	9969	38613	7	4294	549224	0
6184	35243	561	7655	7140	280	101	593104	224
5717	35292	357	7546	7032	255	0	606427	0
21634	32544	121	13118	372	157	4317	593931	0
7771	29430	370	6763	1158	318	3738	594769	4424

Sr

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274
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409
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111
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75
21
183
172
66
330
166

Table S2. Major elements, Be, Cr, Co, Ni, Sr, La and U concentrations (ppm) measured by ICP-MS for 68 sam

Type	Sub-type	Track	Sample name	Be	Na	Mg	Al
S-type	Chondritic	4	IS4 SPH4		9542	177567	4157
S-type	Chondritic	4	IS4 SPH5		4832	172742	16224
S-type	Chondritic	13	IS13 SPH4	0.156	6421	166184	9791
S-type	Chondritic	17	IS17 SPH2		178	133511	15097
S-type	Chondritic	19	S25 IS19 SPH2		1919	140779	7730
S-type	Chondritic	4	S12 IS4D SPH3		184	146583	15371
S-type	Chondritic	4	S9 IS4A SPHS		155	161306	15441
S-type	Chondritic	22	IS22 SPH2		642	155559	11244
S-type	Chondritic	22	IS22 SPH3	3.318	763	164573	9706
S-type	Chondritic	13	(4) IS13 SPH11	0.008	759	177936	5736
S-type	Chondritic	14	(8) IS14 SPH8	0.019	656	155222	14846
S-type	Chondritic	22	(16) IS22 SPH5	0.0004	1163	152596	13280
S-type	Chondritic	4	(20) IS4 SPH9		2195	145125	15516
S-type	Chondritic	8	(23) IS8 SPH5		640	157212	11706
S-type	Chondritic	13	(27) IS13 SPH7		181	139668	17950
S-type	Chondritic	13	(29) IS13 SPH10		348	183043	4311
S-type	non-Chondritic	14	(7) IS14 SPH7		44	216146	412
S-type	Chondritic	14	(9) IS14 SPH9	0.004	376	137838	8850
S-type	Chondritic	22	(1) IS22 SPH1	0.426	8168	104355	22530
I-type	Ni-rich	13	IS13 SPH3		749	235	1748
I-type	Ni-rich	4	S11 IS4C SPH2		176	19263	2512
I-type	Ni-rich	14	(12) IS14 SPH14		664	628	2047
I-type	Ni-rich	8	(24) IS8 SPH6		183	85	927
I-type	Ni-rich	17	17NMAG_37		2680	3197	521
I-type	Ni-poor	8	IS8 SPH1	0.019	1578	964	875
I-type	Ni-poor	8	IS8 SPH2		330	798	1531
I-type	Ni-poor	16	S7 IS16A SPH1		850	166	348
I-type	Ni-poor	12	S16 IS12C SPH1	0.629	623	345	7764
I-type	Ni-poor	19	S23 IS19 SPH1		536	573	404
I-type	Ni-poor	19	S29 IS19 SPH7		822	725	1603
I-type	Ni-poor	8	S5 IS8 SPHR	0.199	384	401	856
I-type	Ni-poor	13	S19 IS13B SPH2		203	295	188
I-type	Ni-poor	22	(2) IS22 SPH4	0.012	387	100	109
I-type	Ni-poor	13	(5) IS13 SPH9	0.008	154	339	151
I-type	Ni-poor	14	(10) IS14 SPH10		514	327	570
I-type	Ni-poor	8	(22) IS8 SPH4		1765	10533	17786
I-type	Ni-poor	13	(26) IS13 SPH6		320	87	367
I-type	Ni-poor	17	17NMAG_15	0.465	5316	6178	10622
I-type	Ni-poor	4	(18) IS4 SPH7	0.010	3943	7496	15125
I-type	Ni-poor	13	(30) IS13 SPH12	0.030	2069	11399	14476
I-type	Ni-poor	14	(31) IS14 SPH5	0.006	2101	8362	13894
I-type	Ni-poor	4	(17) IS4 SPH6	2.62	979	9470	10844

D-type	High Sr, high Si	17	IS17 SPH1	1.46	61972	32998	210325
D-type	High Sr, high Si	17	(14) IS17 SPH4	1.42	61916	28451	209998
D-type	High Sr, high Si	13	(28) IS13 SPH8b	2.03	49383	33244	195530
D-type	High Sr, high Si	17	(13) IS17 SPH3	1.53	49431	27214	122313
D-type	High Sr, high Si	17	(15) IS17 SPH5	1.22	41987	24935	124433
D-type	High Sr, high Si	17	17NMAG_36	0.83	18354	12657	59920
D-type	High Sr, low Si	17	17NMAG_6	2.87	11147	37615	124736
D-type	High Sr, low Si	14	(32) IS14 SPH6	5.05	1712	6220	94992
D-type	High Sr, low Si	14	IS14 SPH2	1.51	13933	7114	57204
D-type	High Sr, low Si	19	19NMAG_50	2.37	910	6387	53827
D-type	High Sr, low Si	19	19NMAG_52	3.13	7914	10729	46119
D-type	Low Sr, low Si	14	(11) IS14 SPH11	6.31	6653	10818	43731
D-type	Low Sr, low Si	8	(21) IS8 SPH3	2.01	1539	20310	35301
D-type	Low Sr, low Si	19	S31 IS19 SPH8		211	3912	84288
D-type	BeLaU low Si	4	S10 IS4B SPH1	14.7	13864	2907	88610
D-type	BeLaU low Si	14	S21 IS14 SPH1	7.8	2618	2231	53878
D-type	BeLaU low Si	14	(6) IS14 SPH4	10.0	7902	4911	90915
D-type	BeLaU low Si	4	(19) IS4 SPH8	32.4	3977	2704	115941
D-type	BeLaU low Si	14	IS14 SPH3	11.4	9001	3336	49658
D-type	BeLaU low Si	17	17NMAG_35	34.6	13854	8615	137992
D-type	BeLaU low Si	17	17NMAG_33	17.6	4740	8348	136612
D-type	BeLaU low Si	17	17NMAG_2	15.0	11780	14941	132262
D-type	BeLaU low Si	19	19MAGx_4	13.5	6623	10360	126494
D-type	BeLaU low Si	17	17NMAG_29	13.4	13895	7152	118488
D-type	BeLaU high Si	13	(25) IS13 SPH5	114.7	13169	8884	236886
D-type	BeLaU high Si	17	17NMAG_5	19.1	17227	9240	148150

ples (data plotted in Figures 9b, 10b, 11, 12, 13, 14, 15).

P	K	Ca	Ti	Cr	Mn	Fe	Co	Ni
546	702	14512	486	2755	1651	201084	326	1484
578	1667	42093	689	762	1838	174601	124	421
569	1310	10107	321	8850	2858	222319	159	6262
1782	581	62464	1139	6359	1359	254052	699	12225
2234	3786	32720	509	6493	1911	271410	578	10610
473	1146	14659	703	4430	2098	269312	699	10253
220	808	19190	678	2109	1946	231199	368	3727
443	3925	21412	934	2742	1953	245727	401	5558
1007	3428	18623	214	3817	1983	221658	630	10550
635	412	3871	479	4363	2098	207312	465	10511
535	343	14309	759	2640	2275	246444	887	12302
1469	75	15121	803	5394	2203	257609	881	4653
126	304	7850	612	4171	2009	280680	462	9650
316			829	3515	1938	265567	534	7504
742		24680	682	4647	2272	278641	652	5677
312	187		343	5819	1459	207997	567	6954
18	53	936	54	49	3043	124011	1.22	
370	220	17121	377	3291	1898	304464	477	6812
7908	3340	174568	14586	6	6351	197902	63	12
88	632	8396	405	5174	4	696801	2486	12023
126	567	3019	92	500	212	618980	4145	56989
	981		189	6455	10	695233	3114	24436
68	43		12	237	14	711965	2828	16474
447	872	10010	38	6278	5361	640609	63	5032
410	318	6818	610	184	2838	708703	7.28	37.3
468	615	14555	269	203	4398	699909	7.56	55.1
117	211	1154	28	209	6381	716287	2.91	6.5
41	370	2509	133	4	11	708875	6.78	32.3
324	607	16015	8	226	4673	700901	17.0	85.6
250	409	15526	2696	382	4324	697265	11.1	90.6
211	261	1797	660	255	2642	715797	61.4	287.2
70	179	590	320	252	4366	714310	7.8	18.0
179	31	2578	493	443	4164	716564	27.0	142.5
43	38		1191	454	3912	715407	57.4	242.4
238	394	6080		77	3636	712442	12.2	50.0
464	229		30503	30	3195	645119	145.5	11.0
52				69	3133	721557	15.3	37
285	7763	35475	64502	1421	77906	336797	21.5	140
5157	1652	7023	37988	37	5396	634286	117.9	31
3282	613	7300	47461	100	5250	620426	113.1	11
1243	727	1170	39657	6	7442	642043	83.2	47
352	221	18309	457	87	14757	651822	19.0	15

3898	27420	140984	5521	87	2492	126468	34.8	133
3892	26583	121006	7407	63	2377	159124	43.2	178
3517	25784	133210	7319	28	3300	162544	49.7	99
2038	22566	304747	5688	19	2582	111080	39.3	25
1878	22661	316450	5765	29	2382	107793	36.5	27
1512	10441	146340	2888	47	1613	53625	21.1	39
2402	6474	33724	1781	102	5551	392982	10.4	138
1552	1220	21112	4442	109	515	534548	36.9	190
34	1716	9908	1673	56	621	601274	25.0	86
293	2275	20653	3303	64	342	605322	12.6	35.37
495	3123	24108	4468	470	291	593238	23.1	33.68
1005	9972	26224	829	40	16407	582408	13.4	33
1899	913	6588	2063	39	12966	598647	14.8	69
390	537	8430	740	223	863	587174	18.3	48.4
449	4759	30786	5005	140	267	543731	38.1	116
982	1350	12378	2729	1307	57	571591	41.7	72
161	2276	15958	5050	765	202	554415	31.3	37
593	2772	2901	2507	82	84	546719	194.3	52
453	7259	15157	3940	138	550	384635	39.1	67
919	25489	16249	7843	321	847	235061	36.7	138
8241	5146	32043	6249	394	6694	322928	102.5	425
3492	6217	36015	9701	243	7383	349500	170.1	842
8185	7479	35109	7580	260	6131	391681	80.7	363
1106	4159	32561	1312	94	112	395260	84.7	259
548	12044	93260	7045	11	550	250278	281.9	196
647	5724	42783	8606	8838	217	192194	64.9	149

Sr	La	U
5.066	0.331	
11.700	0.461	0.055
18.076	0.489	0.076
13.969	0.596	0.002
20.091	0.536	0.022
13.172	0.306	
42.101	0.407	
12.414	0.191	
24.518	0.397	
27.48	0.371	0.016
11.67	0.303	
15.89	0.571	0.033
22.51	0.664	0.003
6.04	0.386	
19.19	0.848	0.027
10.23	0.270	0.054
0.31	0.008	
9.95	0.404	
110	16.520	0.310
2.296	0.182	0.022
3.114	0.008	

103.30	0.24	0.81
57.099	0.339	0.293
6.909	0.326	0.062
1.322	0.026	0.007
11.184	2.001	0.115
4.070	0.122	0.001
13.463	0.188	0.048
8.264	0.588	0.104
1.017	0.026	0.006
5.303	0.090	0.015
1.077	0.024	0.062
2.389		
20.973	0.418	0.101
1.913	0.024	0.002
50.9	4.54	3.14
30.0	13.233	0.135
27.4	3.847	0.091
25.8	2.734	0.119
117.4	5.735	0.696

1349	20.056	1.562
1377.9	19.994	1.373
1435.2	23.869	1.853
2128.1	19.699	1.389
2230.3	19.724	1.584
995.1	9.42	14.58
930.4	20.81	5.39
1172.3	25.67	3.42
859.9	11.160	1.180
508.3	34.976	4.304
651.2	19.294	2.820
119.1	5.890	0.998
132.8	15.998	2.147
3.730	0.214	0.109
447.1	66.3	5.114
2242.8	159.4	6.148
2021.6	81.2	7.685
128.6	76.4	2.595
243.2	31.45	2.98
307.7	107.25	8.69
1071.0	82.99	15.17
962.6	52.55	17.31
1517.1	51.86	10.30
577.9	120.34	2.54
1792.7	260.4	9.715
2892	71.97	11.7

Table S3. Elemental data (ppm) measured by ICP-MS for 12 BeLaU-type spherules and the CI normalizing val

Type	Sub-type	Track	Sample name	Li	Be	Na	Mg
D-type	BeLaU low Si	4	S10 IS4B SPH1	84.7	14.7	13864	2907
D-type	BeLaU low Si	14	S21 IS14 SPH1	62.2	7.8	2618	2231
D-type	BeLaU low Si	14	(6) IS14 SPH4	87.0	10.0	7902	4911
D-type	BeLaU low Si	4	(19) IS4 SPH8	22.4	32.4	3977	2704
D-type	BeLaU low Si	14	IS14 SPH3	37.7	11.4	9001	3336
D-type	BeLaU low Si	17	17NMAG_35	162.7	34.6	13854	8615
D-type	BeLaU low Si	17	17NMAG_33	108.9	17.6	4740	8348
D-type	BeLaU low Si	17	17NMAG_2	170.0	15.0	11780	14941
D-type	BeLaU low Si	19	19MAGx_4	216.0	13.5	6623	10360
D-type	BeLaU low Si	17	17NMAG_29	102.2	13.4	13895	7152
D-type	BeLaU high Si	13	(25) IS13 SPH5	120.4	114.7	13169	8884
D-type	BeLaU high Si	17	17NMAG_5	154.4	19.1	17227	9240
CI(ppm) Anders and Grevesse (1989)				1.5	0.025	5000	98990

ues of Anders and Grevesse (1989).

Al	Si	P	K	Ca	Sc	Ti	V	Cr
88610		449	4759	30786	36.3	5005	171	140
53878		982	1350	12378	20.9	2729	101	1307
90915		161	2276	15958	43.2	5050	220	765
115941		593	2772	2901	14.4	2507	57	82
49658		453	7259	15157	11.2	3940	98	138
137992		919	25489	16249	27.2	7843	202	321
136612		8241	5146	32043	34.0	6249	465	394
132262		3492	6217	36015	51.9	9701	277	243
126494		8185	7479	35109	31.9	7580	266	260
118488		1106	4159	32561	18.8	1312	36	94
236886		548	12044	93260	79.3	7045	323	11
148150		647	5724	42783	80.3	8606	372	8838
8680	106400	1220	558	9280	5.82	436	56.5	2660

Mn	Fe	Co	Ni	Cu	Zn	Ga	Ge	As
267	543731	38	116	117	44	42.3	8.44	2.06
57	571591	42	72	54	9	27.7	9.50	1.04
202	554415	31	37	47		20.2	9.21	1.41
84	546719	194	52	125		61.2	9.11	13.06
550	384635	39	67		1414	14.6	6.53	1.00
847	235061	37	138		478	28.8	3.91	2.96
6694	322928	103	425	24	542	51.6	5.57	2.59
7383	349500	170	842		1314	73.0	6.50	4.47
6131	391681	81	363	346	83	103.6	6.79	3.77
112	395260	85	259		1503	24.5	6.96	1.82
550	250278	282	196	108		68.8	4.90	4.24
217	192194	65	149	98	97	89.5	3.51	1.70
1990	190400	502	11000	126	312	10	32.7	1.86

Se	Rb	Sr	Y	Zr	Nb	Mo	Ag	Cd
0.25	17.6	447	135	465	38.2	4.9	0.24	0.09
0.25	4.5	2243	120	247	15.1	77.5	0.07	0.05
0.24	6.8	2022	133	626	34.5	19.8	0.21	0.06
0.24	13.9	129	188	199	7.1	22.7	0.11	
0.13	40.9	243	44	131	12.8	6.7	0.05	0.29
0.28	173.9	308	86	441	29.2	10.3		0.03
0.11	27.8	1071	63	477	34.0	18.2	0.16	0.08
0.19	31.7	963	77	423	23.2	13.4	0.15	1.59
0.16	33.7	1517	61	420	21.0	14.2	0.08	0.12
0.37	11.6	578	255	153	4.3	27.4	0.03	0.02
0.71	46.7	1793	487	1008	84.8	179.2	0.23	0.24
0.44	22.6	2892	214	1117	71.2	5.4	0.304	0.132
18.6	2.3	7.8	1.56	3.94	0.246	0.928	0.199	0.686

Sb	Cs	Ba	La	Ce	Pr	Nd	Sm	Eu
1.28	1.09	914	66.3	128	14.5	58.2	12.4	2.68
0.05	0.18	759	159.4	267	25.6	89.4	16.0	2.96
0.55	0.37	1452	81.2	152	16.7	68.0	15.6	3.35
0.68	1.92	228	76.4	181	20.3	85.5	18.6	3.20
2.83	2.89	385	31.5	66	7.1	27.1	6.8	1.18
12.41	17.80	792	107.2	227	24.6	100.5	11.2	3.69
9.70	4.94	1404	83.0	214	25.7	99.4	14.0	2.57
11.61	4.62	1613	52.6	116	14.1	60.2	11.1	2.35
1.96	3.17	3052	51.9	110	12.4	50.3	11.3	2.57
10.35	2.27	740	120.3	170	20.5	88.7	22.8	4.83
2.35	5.67	2848	260.4	451	49.8	204.6	45.5	8.87
0.94	2.68	2573	72.0	159	19.2	82.5	21.2	4.90
0.142	0.187	2.34	0.235	0.603	0.089	0.452	0.147	0.056

Gd	Tb	Dy	Ho	Er	Tm	Yb	Lu	Hf
15.30	2.68	17.36	3.79	11.26	1.80	10.57	1.58	10.92
17.87	2.72	17.26	3.70	11.59	1.85	11.92	1.79	5.95
17.23	2.93	19.56	4.12	13.81	2.16	14.57	2.08	14.27
19.96	3.37	23.51	5.08	17.00	2.70	16.00	2.24	6.22
5.61	1.06	5.53	1.21	3.30	0.59	3.55	0.51	3.45
13.33	2.23	13.23	2.32	8.18	0.96	6.98	1.78	7.97
11.07	1.98	11.37	2.25	7.30	1.05	6.96	1.31	10.88
11.83	2.14	12.42	2.81	7.02	1.34	7.81	1.49	9.93
10.53	1.56	10.20	2.08	5.80	1.03	5.64	0.83	10.90
29.60	5.14	34.05	8.09	23.62	3.85	22.88	4.00	3.65
51.72	9.28	64.72	15.28	55.64	9.26	65.66	10.34	24.11
23.30	4.49	29.94	6.85	20.94	3.47	22.72	3.45	23.20
0.197	0.0363	0.243	0.0556	0.159	0.0242	0.163	0.0243	0.104

Ta	W	Re	Tl	Pb	Bi	Th	U
1.45	17.96	0.001	0.322	3.76	0.032	14.70	5.11
0.63	7.53	0.002		0.34	0.002	8.41	6.15
1.34	4.39	0.002	0.001	1.01		15.32	7.68
0.60	8.62	0.002		1.55		8.26	2.60
0.76	7.12	0.131	0.338	2.20	0.461	9.90	2.98
1.50	12.64	0.015	1.221	4.35	0.999	38.73	8.69
1.23	20.12	0.022	0.611	17.85	0.119	19.80	15.17
1.40	9.28	0.010	0.153	10.05	0.656	56.05	17.31
1.57	12.13	0.002	0.010	2.46	0.079	23.33	10.30
0.08	669.18	0.015	0.210	3.90	0.557	4.13	2.54
2.64	63.34	0.008	0.044	10.49	0.021	18.76	9.72
2.03	6.71	0.005	0.053	7.78	0.040	24.28	11.69
0.014	0.093	0.0365	0.142	2.47	0.114	0.0294	0.0081